Introduction to Nuclear Fusion

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To build a sun on earth



What is closed magnetic confinement?

Open Magnetic System



- Suffering from end losses

J.P. Freidberg, "Ideal Magneto-Hydro-Dynamics", lecture note A. A. Harms et al, "Principles of Fusion Energy", World Scientific (2000)

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ion





Toroidal Field (TF) coil



Applying toroidal magnetic field 3.5 T in KSTAR, 5.3 T in ITER



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Magnetic field of earth? 0.5 Gauss = 0.00005 T

http://www.transformacionconciencia.com/archives/2384 13







What kind of drift motions?

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Where does the gradient come from?





 $\nabla \times \mathbf{B} = \mu_0 \mathbf{J}$ $\oint \mathbf{B}_{\phi} \cdot d\mathbf{l} = \mu_0 N I_c$ $B_{\phi}(R) = \frac{\mu_0 N I_c}{2\pi R}$



$$\mathbf{v}_{D,R} = \frac{mv_{\parallel}^2}{qB_0^2} \frac{\mathbf{R}_0 \times \mathbf{B}_0}{R^2}$$
$$\mathbf{v}_{D,\nabla B} = \pm \frac{1}{2} v_{\perp} r_L \frac{\mathbf{B} \times \nabla B}{B^2}$$
$$= \frac{mv_{\perp}^2}{2qB} \frac{\mathbf{B} \times \nabla B}{B^2}$$

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$$\mathbf{v}_D = \frac{m}{q} \frac{1}{R_0 B_{\phi}(R_0)} \left[v_{\parallel}^2 + \frac{v_{\perp}^2}{2} \right] \mathbf{e}_Z$$

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Electric field, **E**





Poloidal magnetic field required Tokamak .VS. Stellarator

What is a tokamak?



Is the plasma current driven if toroidal magnetic filed is applied?

Poloidal magnetic field required How to drive plasma current?

Tokamak





Tokamak



Pulsed Operation



Steady-State Operation





Adding vertical (equilibrium) field coils (PF: Poloidal Field)





KSTAR



Adding vertical (equilibrium) field coils (PF: Poloidal Field)



Tokamak



Force balance by vertical field coils: Plasma positioning





The plasma shape can be modified by PF coil currents.





Adding vertical (equilibrium) field coils (PF: Poloidal Field)

Plasma positioning & shaping by PF coils



Invented by Igor Tamm and Andrei Sakharov in 1952









Toroidalnaja kamera magnitnaja katushka (Toroidal chamber magnetic coil)




Cutaway of the Toroidal Chamber in Artsimovitch's Paper Research on Controlled Nuclear Fusion in the USSR





WIKIPEDIA

Toroidalnaja kamera magnitnaja katushka (Toroidal chamber magnetic coil)

1958 IAEA FEC, Geneva, Switzerland



T1: The world's first tokamak, Kurchatov Institute, Moscow Russia

JET (Joint European Torus): $R_0 = 3 \text{ m}$, a = 0.9 m, 1983-today





JET (Joint European Torus): $R_0 = 3 \text{ m}$, a = 0.9 m, 1983-today





JET (Joint European Torus): $R_0 = 3 \text{ m}$, a = 0.9 m, 1983-today





http://cafe.naver.com/gamebox26.cafe?iframe_url=/ArticleRead.nhn%3Farticleid=16140 42

















KSTAR (Korea Superconducting Tokamak Advanced Research): $R_0 = 1.8 \text{ m}, a = 0.5 \text{ m}, 2007 \text{-today}$

KSTAR 1st plasma

Analyse the KSTAR 1st plasma



How to characterize a tokamak?



Cylindrical and local coordinates for a tokamak



ex) KSTAR: 3.6, ITER: 3.1

• Plasma equilibrium parameters



- Elongation: κ
- Triangularity: δ

• Plasma equilibrium parameters



$$\kappa = \frac{b}{a}$$
 $\delta = \frac{c+d}{2a}$

• Separation of plasma from wall by a limiter and a divertor



• Plasma equilibrium parameters

Parameters	KSTAR	ITER	- Plasma shape
Major Radius, R_0 Minor Radius, a Plasma Current, I_P Elongation, κ_x Triangularity, δ_x Toroidal Field, B_0 Pulse LengthFuel	1.8 m 0.5 m 2.0 MA 2.0 0.8 3.5 T 300 s H, D	6.2 m 2.0 m 15 MA 1.85 0.5 5.3 T 500 s D, T	JET JT-60U KSTAR Asdex-U DIII-D

1m

Magnetic Flux Surface

- In a tokamak configurations with confined plasmas the magnetic lines lie on a set of nested toroidal surfaces called flux surfaces.
- Pressure is constant along a magnetic field line.
- Magnetic lines lie in surfaces of constant pressure.
- Flux surfaces are surfaces of constant pressure.
- The current lines lie on surfaces of constant pressure.



$$\vec{J} \times \vec{B} = \nabla p$$
$$\vec{B} \cdot \nabla p = 0 \quad \vec{J} \cdot \nabla p = 0$$

Magnetic Flux Surface









http://www.freewebs.com/weatherexplorer/apps/blog/show/286262 http://www.econym.demon.co.uk/isotut/isobars.htm

• Safety factor q = number of toroidal orbits per poloidal orbit



• Safety factor q = number of toroidal orbits per poloidal orbit



Magnetic Shear

- Measuring the change in pitch angle of a magnetic field line from one flux surface to the next
- Playing an important role in stabilizing MHD instabilities, particularly those driven by the pressure gradient:

A perturbation aligned with $\mathbf{B}(r)$ will, at a point with increased minor radial distance r+dr, encounter field lines at a different angle which again will vary as the perturbation grows to another distance r+dr'. Any helically resonant instabilities are thus radially localised.



 $s(r) \equiv \frac{r}{q} \frac{dq}{dr}$

$$\nabla p = \vec{J} \times \vec{B}$$
$$\nabla \times \vec{B} = \mu_0 \vec{J}$$
$$\nabla \cdot \vec{B} = 0$$

- Consider the axisymmetric torus, the simplest, multi-dimensional configuration
- We shall derive the Grad–Shafranov equation for axisymmetric equilibria.
- This provides a complete description of toroidal equilibrium: radial pressure balance, toroidal force balance,
 β limits, q-profiles, etc.

The Grad-Shafranov Equation

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - F \frac{dF}{d\psi}$$

- Nonlinear
- Partial differential equation
- Grad and Rubin (1958), Shafranov (1960)
- Toroidal axisymmetric $\partial/\partial \Phi = 0$





The Grad-Shafranov Equation

Sequence of solution of the MHD equilibrium equations 1. The $\nabla \cdot \mathbf{B} = 0$

- 2. Ampere's law: $\mu_0 \mathbf{J} = \nabla \mathbf{x} \mathbf{B}$
- 3. The momentum equation: $JxB = \nabla p$

$$\nabla p = \vec{J} \times \vec{B}$$
$$\nabla \times \vec{B} = \mu_0 \vec{J}$$
$$\nabla \cdot \vec{B} = 0$$

$$J \times B = \nabla p$$
$$\nabla p = \left(\vec{J}_{\phi} + \vec{J}_{p}\right) \times \left(\vec{B}_{\phi} + \vec{B}_{p}\right)$$

The Grad-Shafranov Equation

• Momentum equation $\vec{J} \times \vec{B} = \nabla p$

 $\nabla p = \left(\vec{J}_{\phi} + \vec{J}_{p} \right) \times \left(\vec{B}_{\phi} + \vec{B}_{p} \right)$

- aim: to express each term with ψ and F

$$\psi = \frac{\psi_p}{2\pi}, \quad F = \frac{\mu_0 I_p}{2\pi}$$

$$\nabla \cdot \vec{B} = 0 \qquad \longrightarrow \quad \vec{B}_p = -\frac{1}{R} \vec{e}_{\phi} \times \nabla \psi$$

$$\mu_0 \vec{J}_p = \nabla \times \vec{B}_\phi \quad \longrightarrow \quad \vec{J}_p = -\frac{1}{\mu_0 R} \vec{e}_\phi \times \nabla F, \quad B_\phi = \frac{F}{R}$$

$$\mu_0 \vec{J}_{\phi} = \left(\nabla \times \vec{B} \right)_{\phi} \longrightarrow J_{\phi} = -\frac{1}{\mu_0 R} \Delta^* \psi$$

- The Grad-Shafranov Equation
 - The \bigtriangledown ·**B** Equation

In cylindrical coordinates for toroidal axisymmetric fields $(\partial/\partial \Phi = 0)$

Ψ: Stream function for the poloidal magnetic field

The Grad-Shafranov Equation

- The stream function ψ is closely related to the poloidal flux in the plasma.

$$\begin{split} \nu_{p} &= \int \vec{B}_{p} d\vec{A} \\ &= \int_{0}^{2\pi} d\phi \int_{R_{a}}^{R} R dR B_{Z}(R, Z = 0) \\ &= \int_{R_{a}}^{R} 2\pi R \frac{1}{R} \frac{\partial \psi}{\partial R} dR \quad \longleftarrow \quad B_{Z} = \frac{1}{R} \frac{\partial \psi}{\partial R} \\ &= 2\pi [\psi(R, 0) - \psi(R_{a}, 0)] = 2\pi \psi \end{split}$$

Poloidal flux on axis is zero

The Grad-Shafranov Equation

- The \bigtriangledown ·**B** Equation

$$\begin{aligned} \nabla \cdot \vec{B} &= 0 & \frac{1}{R} \frac{\partial (RB_R)}{\partial R} + \frac{1}{R} \frac{\partial B_{\phi}}{\partial \phi} + \frac{\partial B_Z}{\partial Z} = 0 \\ \vec{B} \cdot \nabla \psi &= 0 & B_R \frac{\partial \psi}{\partial R} + B_Z \frac{\partial \psi}{\partial Z} = 0 & \psi: \text{ Stream function for the poloidal magnetic field} \\ \text{Magnetic flux surface} & B_R = -\frac{1}{R} \frac{\partial \psi}{\partial Z}, \quad B_Z = \frac{1}{R} \frac{\partial \psi}{\partial R} & \text{ for the poloidal magnetic field} \\ \vec{B} &= B_{\phi} \vec{e}_{\phi} + \vec{B}_{p} & 2\pi d\psi = 2\pi (\psi + d\psi) - 2\pi \psi = \int_{A + \partial A} \vec{B}_p \cdot d\vec{A} - \int_A \vec{B}_p \cdot d\vec{A} = \int_{\partial A} \vec{B}_p \cdot d\vec{A} \approx B_p 2\pi R dr \\ RB_p &= \frac{d\psi}{dr} = |\nabla \psi| & \longleftarrow 2\pi \psi = \psi_p = \int \vec{B}_p \ d\vec{A} & \vec{B}_p = -\frac{1}{R} \vec{e}_{\phi} \times \nabla \psi \end{aligned}$$

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The Grad-Shafranov Equation

- Ampere's law

$$\begin{split} \mu_{0}\vec{J}_{p} &= \nabla \times \vec{B}_{\phi} \\ \int_{\delta A}^{\nabla} \times \vec{B}_{\phi} \cdot d\vec{A} = \oint \vec{B}_{\phi} \cdot d\vec{l} = 2\pi d(RB_{\phi}) = 2\pi d\psi \frac{\partial(RB_{\phi})}{\partial\psi} \\ \int_{\delta A} \mu_{0}\vec{J}_{p} \cdot d\vec{A} &= \mu_{0}J_{p}2\pi R dr \\ \Rightarrow \frac{\partial(RB_{\phi})}{\partial\psi} &= \frac{\mu_{0}J_{p}R}{(d\psi/dr)} \\ \Rightarrow \mu_{0}J_{p}R &= \frac{\partial F}{\partial\psi} \frac{\partial \psi}{\partial r} = \frac{\partial F}{\partial\psi} |\nabla\psi| = \frac{\partial F}{\partial r} = |\nabla F(\psi)| \quad \longleftarrow \quad F(\psi) \equiv RB_{\phi} \\ \vec{J}_{p} &= -\frac{1}{\mu_{0}R} \vec{e}_{\phi} \times \nabla F, \quad B_{\phi} = \frac{F}{R} \end{split}$$

The Grad-Shafranov Equation

- Interpretation of $F(\psi)$

$$\begin{split} I_{p} &= \int \vec{J}_{p} \cdot d\vec{A} \\ &= \int_{0}^{2\pi} d\phi \int_{0}^{R} R dR J_{Z}(R, Z = 0) \\ &= \int_{0}^{2\pi} d\phi \int_{0}^{R} R dR \frac{1}{\mu_{0}R} \frac{\partial}{\partial R} (RB_{\phi}) \\ &= \int_{0}^{R} 2\pi R \frac{1}{\mu_{0}R} \frac{\partial F}{\partial R} dR \qquad \longleftrightarrow \qquad J_{Z} = \frac{1}{\mu_{0}} \frac{1}{R} \frac{\partial}{\partial R} (RB_{\phi}), \quad F(\psi) \equiv RB_{\phi} \\ &= 2\pi [F(R,0) - F(0,0)]/\mu_{0} \qquad R = 0 \rightarrow F(\psi) = 0 \\ &= 2\pi F(\psi)/\mu_{0} \end{split}$$

 $I_{\rho}(\psi)$ is the total poloidal current and the toroidal field coils passing through the circle $\psi(R, 0) = \text{const.}$

- The Grad-Shafranov Equation
 - Ampere's law

$$\mu_{0}\vec{J} = \nabla \times \vec{B}$$

$$J_{\phi} = \frac{1}{\mu_{0}} \left(\frac{\partial B_{R}}{\partial Z} - \frac{\partial B_{Z}}{\partial R} \right) = -\frac{1}{\mu_{0}} \left(\frac{\partial}{\partial R} \left(\frac{1}{R} \frac{\partial \psi}{\partial R} \right) + \frac{1}{R} \frac{\partial^{2} \psi}{\partial Z^{2}} \right)$$

$$\equiv -\frac{1}{\mu_{0}R} \Delta^{*} \psi \qquad \longleftarrow \quad B_{R} = -\frac{1}{R} \frac{\partial \psi}{\partial Z}, \quad B_{Z} = \frac{1}{R} \frac{\partial \psi}{\partial R}$$

$$\Delta^{*} \psi \equiv R \frac{\partial}{\partial R} \left(\frac{1}{R} \frac{\partial \psi}{\partial R} \right) + \frac{\partial^{2} \psi}{\partial Z^{2}} \quad \text{elliptic operator}$$

The Grad-Shafranov Equation

 $\vec{J}_{p} = -\frac{1}{\mu_{0}R}\vec{e}_{\phi} \times \nabla F$ Symmetry $\vec{B}_{p} = -\frac{1}{R}\vec{e}_{\phi} \times \nabla \psi$ - Momentum equation $\vec{J} \times \vec{B} = \nabla p$ $\nabla p = \left(\vec{J}_{\phi} + \vec{J}_{p}\right) \times \left(\vec{B}_{\phi} + \vec{B}_{p}\right)$ $= \left(\vec{J}_{\phi} - \frac{1}{\mu_{0}R}\vec{e}_{\phi} \times \nabla F\right) \times \left(\vec{B}_{\phi} - \frac{1}{R}\vec{e}_{\phi} \times \nabla \psi\right)$ $B_{\phi} = \frac{F(\psi)}{R}$ $= -\vec{J}_{\phi} \times \left(\frac{1}{R}\vec{e}_{\phi} \times \nabla \psi\right) - \left(\frac{1}{\mu_{0}R}\vec{e}_{\phi} \times \nabla F\right) \times \vec{B}_{\phi}$ $J_{\phi} = -\frac{1}{\mu_{0}R} \Delta^{*} \psi$ $= \frac{\vec{J}_{\phi}}{R} \nabla \psi - \frac{\vec{B}_{\phi}}{\mu_{o}R} \nabla F \iff A \times (B \times C) = B(A \cdot C) - C(A \cdot B)$

$$J_{\phi} = R \frac{dp}{d\psi} + \frac{B_{\phi}}{\mu_0} \frac{dF}{d\psi} = R \frac{dp}{d\psi} + \frac{F(\psi)}{\mu_0 R} \frac{dF}{d\psi} = -\frac{1}{\mu_0 R} \Delta^* \psi$$
Tokamak Equilibrium

The Grad-Shafranov Equation



- BCs: provided by the transformer-induced poloidal magnetic field outside the plasma
- In practice, the G-S equation is solved numerically to find the geometrical location of the magnetic surfaces and the radial distribution of the axial current density in a way that is consistent with the experimentally measured pressure profiles (*p*) and the externally applied field (*F*).







Lesch, Astrophysics, IPP Summer School (2008)