

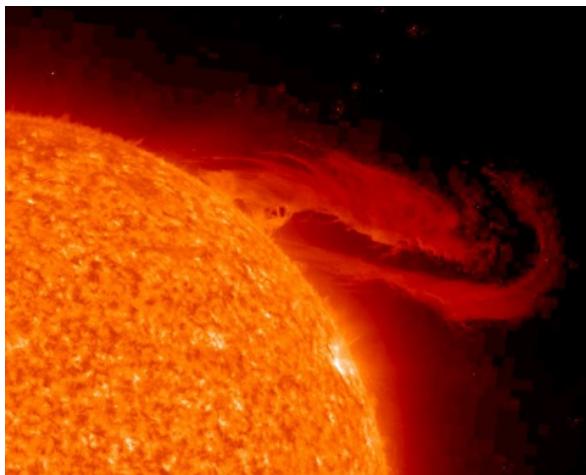
Introduction to Nuclear Fusion

Prof. Dr. Yong-Su Na

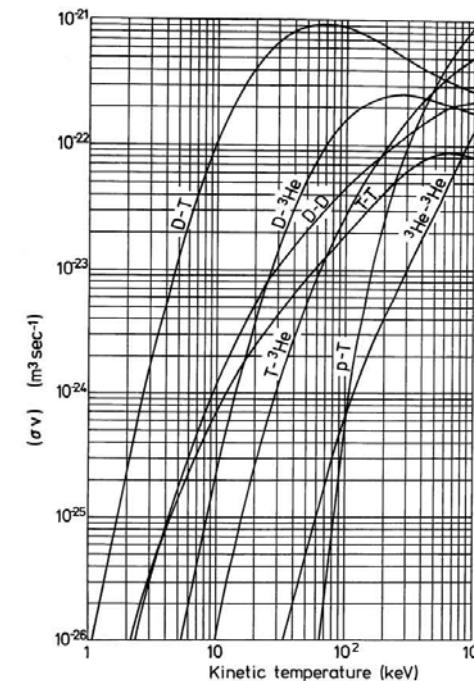
How to describe a plasma?

Description of a Plasma

- Thermonuclear fusion with high energy particles of 10-20 keV energy → plasma



- Three approaches to describe a plasma
 - Single particle approach
 - Kinetic theory
 - Fluid theory



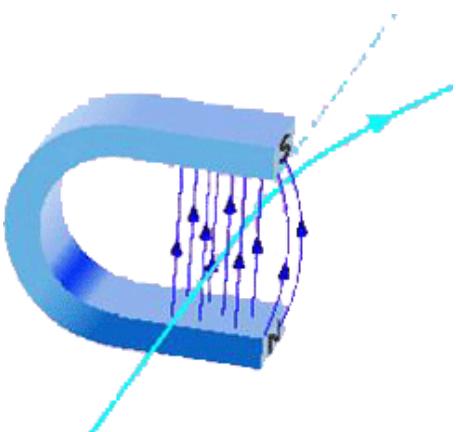
<http://dz-dev.net/blog/tag/eruption-solaire>

Individual Charge Trajectories

- **Equation of motion**

- Basic relation determining the motion of an individual charged particle of mass m and charge q in a combined electric (**E**) and magnetic (**B**) field
- Neglecting electromagnetic fields generated by movement of the charge itself

$$m \frac{d\mathbf{v}}{dt} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$



Individual Charge Trajectories

- Homogeneous electric field

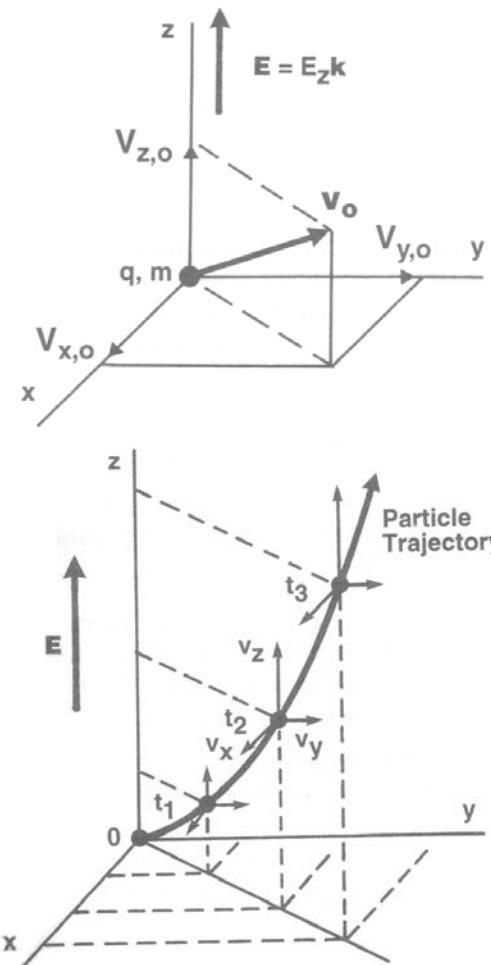
$$m \frac{d\mathbf{v}}{dt} = q\mathbf{E} = qE_z\mathbf{k}$$

$$v_x(0) = v_{x,0}, \quad v_y(0) = v_{y,0}, \quad v_z(0) = v_{z,0}$$

$$v_x(t) = v_{x,0}, \quad v_y(t) = v_{y,0}, \quad v_z(t) = v_{z,0} + \frac{q}{m} E_z t$$

$$x(t) = v_{x,0}t, \quad y(t) = v_{y,0}t,$$

$$z(t) = v_{z,0}t + \frac{1}{2} \left(\frac{q}{m} E_z \right) t^2$$



Individual Charge Trajectories

- Homogeneous magnetic field

$$m \frac{d\mathbf{v}}{dt} = q(\mathbf{v} \times \mathbf{B}) = q(\mathbf{v} \times B_z \mathbf{k})$$

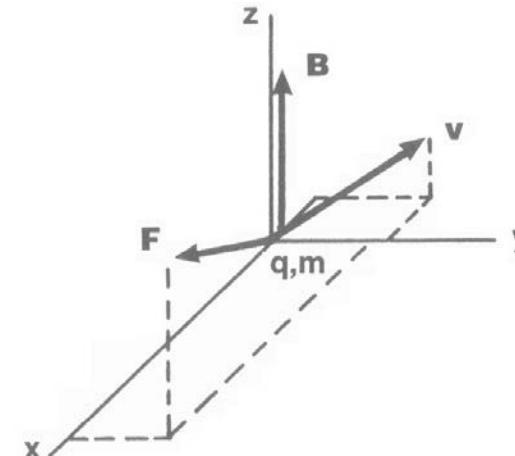
$$v_x(0) = v_{x,0}, \quad v_y(0) = v_{y,0}, \quad v_z(0) = v_{z,0} = v_{\parallel}$$

$$v_x(t) = v_{\perp} \cos(\omega_c t + \phi), \quad v_y(t) = \mp v_{\perp} \sin(\omega_c t + \phi), \quad v_z(t) = v_{z,0} = v_{\parallel}$$

$$v_{\perp} = \sqrt{v_x^2 + v_y^2}, \quad \tan(\phi) = \mp \frac{v_{y,0}}{v_{x,0}}, \quad \omega_c = \frac{|q|B_z}{m} \quad \text{Cyclotron frequency}$$

$$x(t) = x_0 + \frac{v_{\perp}}{\omega_c} \sin(\omega_c t + \phi), \quad y(t) = y_0 \pm \frac{v_{\perp}}{\omega_c} \cos(\omega_c t + \phi), \quad z(t) = z_0 + v_{\parallel} t$$

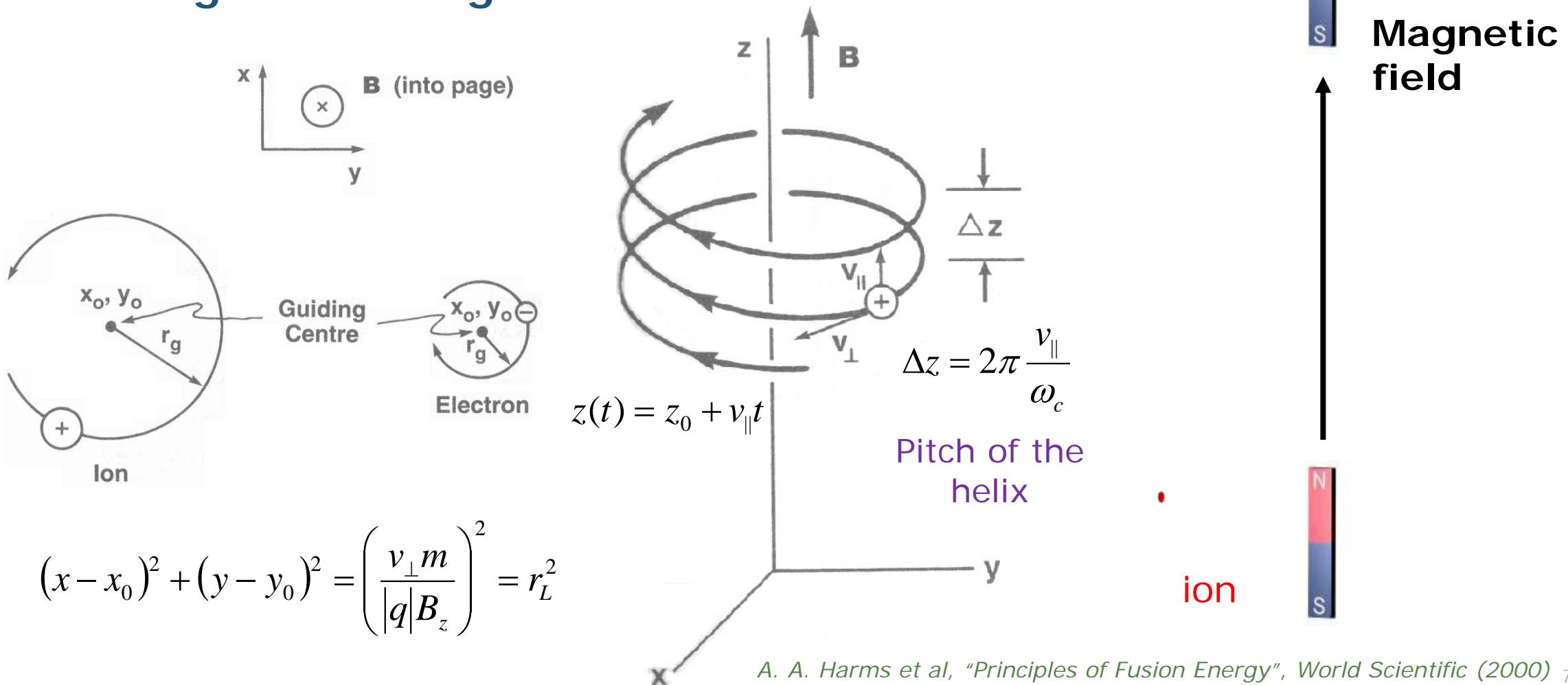
$$(x - x_0)^2 + (y - y_0)^2 = \left(\frac{v_{\perp}}{\omega_c} \right)^2 = \left(\frac{v_{\perp} m}{|q| B_z} \right)^2 = r_L^2 \quad \text{Larmor radius}$$



5 T, 10 keV
 $r_L = 0.05 \text{ mm}$ for e
 $r_L = 2.9 \text{ mm}$ for d
 $r_L = 3.5 \text{ mm}$ for t

Individual Charge Trajectories

- Homogeneous magnetic field



Individual Charge Trajectories

- Combined homogeneous electric and magnetic field

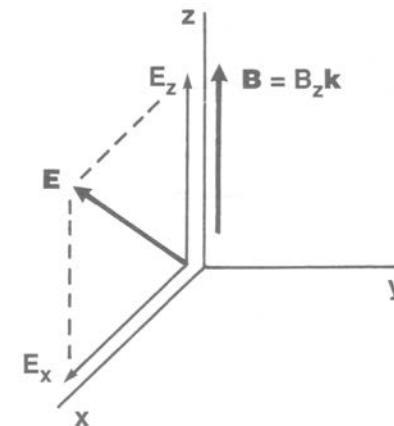
$$m \frac{d\mathbf{v}}{dt} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}) = q[(E_x \mathbf{i} + E_z \mathbf{k}) + \mathbf{v} \times B_z \mathbf{k}]$$

$$v_x(0) = v_{x,0}, \quad v_y(0) = v_{y,0}, \quad v_z(0) = v_{z,0} = v_{\parallel}$$

$$v_x(t) = v_{\perp}^* \cos(\omega_c t + \phi^*), \quad v_y(t) = \mp v_{\perp}^* \sin(\omega_c t + \phi^*) - \frac{E_x}{B_z},$$

$$v_z(t) = v_{\parallel} + \left(\frac{qE_z}{m} \right) t$$

$$v_{\perp}^* = \sqrt{v_x^2 + \left(v_y + \frac{E_x}{B_z} \right)^2} = \frac{v_{\perp} \cos \phi}{\cos \phi^*}, \quad \tan(\phi^*) = \mp \frac{v_{y,0} + \frac{E_x}{B_z}}{v_{x,0}}$$



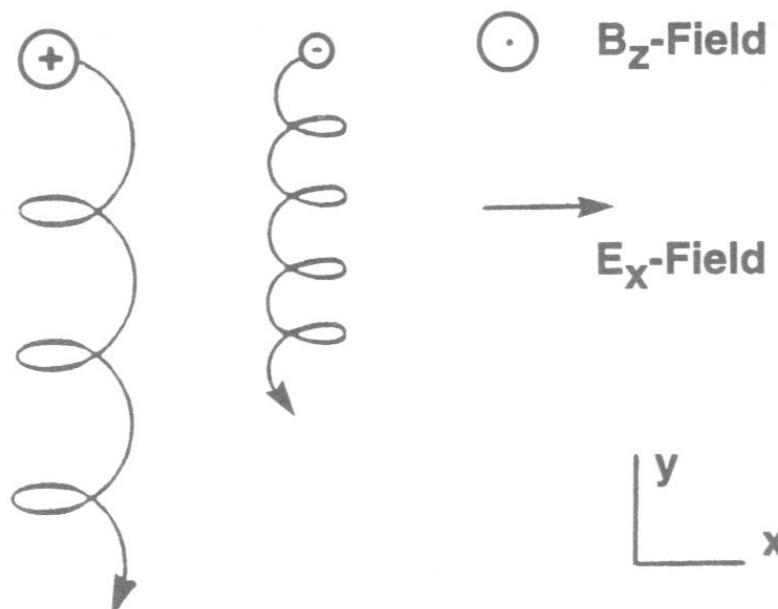
Individual Charge Trajectories

- Combined homogeneous electric and magnetic field

$$x(t) = x_0 + \frac{v_{\perp}^*}{\omega_c} \sin(\omega_c t + \phi^*), \quad y(t) = y_0 \pm \frac{v_{\perp}^*}{\omega_c} \cos(\omega_c t + \phi^*) - \frac{E_x}{B_z} t,$$

$$z(t) = z_0 + v_{\parallel} t + \frac{1}{2} \left(\frac{qE_z}{m} \right) t^2$$

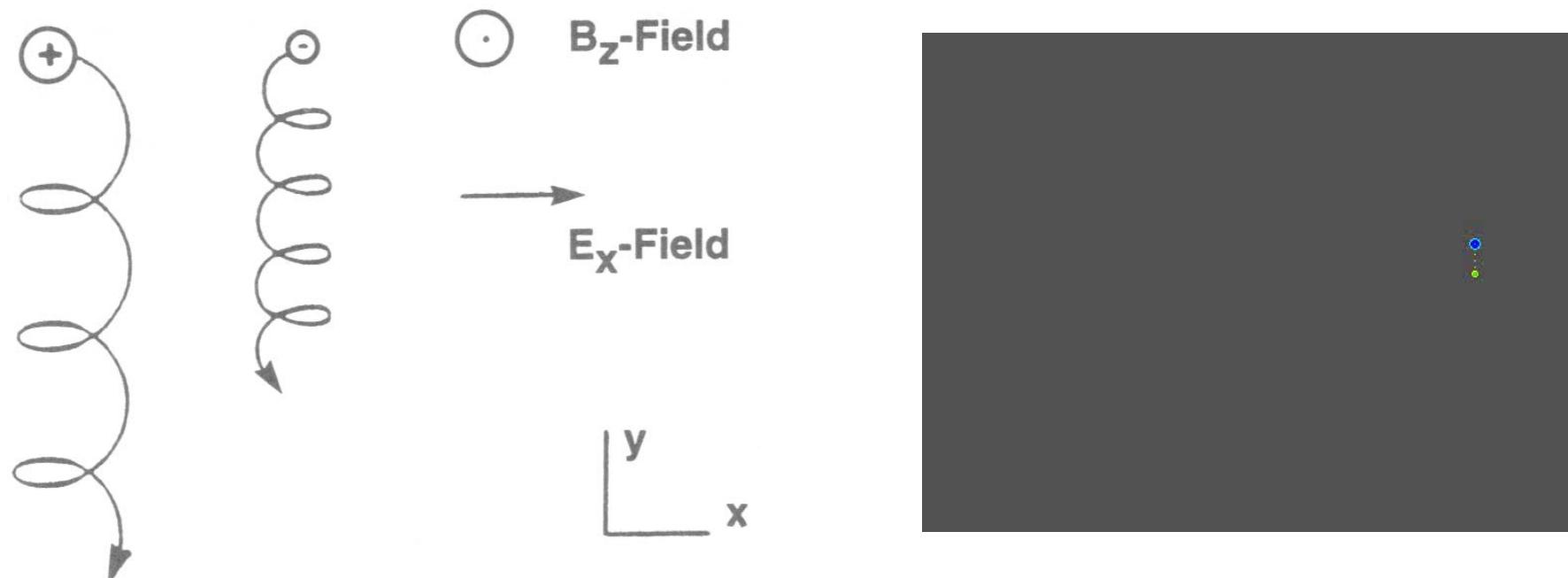
$$(x - x_0)^2 + \left(y - y_0 + \frac{E_x}{B_z} t \right)^2 = \left(\frac{v_0^*}{\omega_c} \right)^2$$



Individual Charge Trajectories

- Combined homogeneous electric and magnetic field

$$\mathbf{v} = \mathbf{v}_{gc} + \mathbf{v}_g \quad : \text{Guiding centre} + \text{Gyro motion}$$



Individual Charge Trajectories

- Combined homogeneous electric and magnetic field

$$\mathbf{v} = \mathbf{v}_{gc} + \mathbf{v}_g \quad : \text{Guiding centre + Gyro motion}$$

2.2

IN MAGNETIC FIELDS

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Because of this change of momentum, the *centre of curvature* of the projection on the xy -plane of the path is displaced a distance

$$\mathbf{D} = \mathbf{p}' - \mathbf{p} \quad (9)$$

or, because of (6)–(8),

$$\mathbf{D} = -\frac{c}{eH^2} [\mathbf{H} \Delta \mathbf{p}_\perp], \quad (10)$$

$$\mathbf{D} = -\frac{c}{eH^2} \left[\mathbf{H} \int \mathbf{f}_\perp dt \right]. \quad (11)$$

This formula holds for a single collision, but of course also for a series of collisions. If \mathbf{f} is a continuous force, \mathbf{D} gives the displacement of that point where the centre of curvature would be if \mathbf{f} vanished for a moment. This point shall be called the guiding centre. If \mathbf{f} is continuous, the guiding centre drifts with the velocity

$$\mathbf{U}_\perp = \frac{d\mathbf{D}}{dt} = -\frac{c}{eH^2} [\mathbf{H} \mathbf{f}]. \quad (12)$$

H. Alfvén, *Cosmical Electrodynamics*, Oxford (1950)



Hannes Alfvén
(1908-1995)
“Nobel prize in
Physics (1970)”



Individual Charge Trajectories

- Combined homogeneous electric and magnetic field

$$\mathbf{v} = \mathbf{v}_{gc} + \mathbf{v}_g \quad : \text{Guiding centre} + \text{Gyro motion}$$

$$m \frac{d\mathbf{v}}{dt} = \mathbf{F} + q(\mathbf{v} \times \mathbf{B})$$

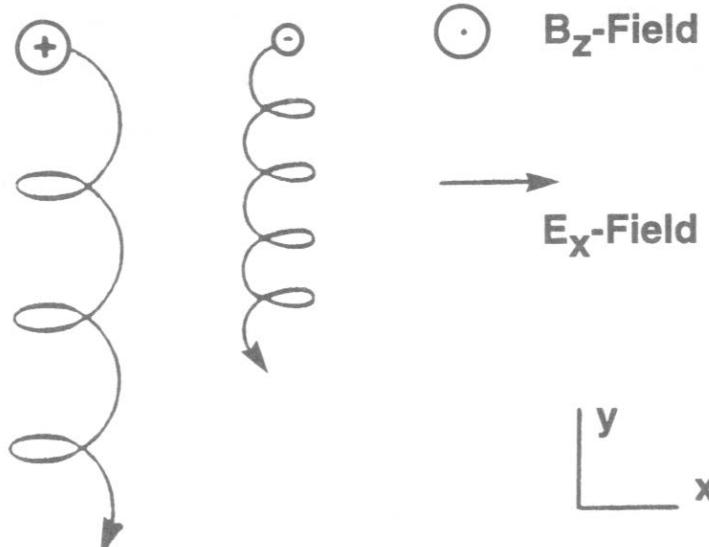
$$m \frac{d\mathbf{v}_{gc}}{dt} + m \frac{d\mathbf{v}_g}{dt} = \mathbf{F} + q(\mathbf{v}_{gc} \times \mathbf{B}) + q(\mathbf{v}_g \times \mathbf{B})$$

$$v_{gc,\parallel}(t) = v_{gc,\parallel}(0) + \frac{1}{m} \int F_{\parallel} dt$$

$$\mathbf{F} = \mathbf{E}q$$

$$\mathbf{v}_{D,E} = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$$

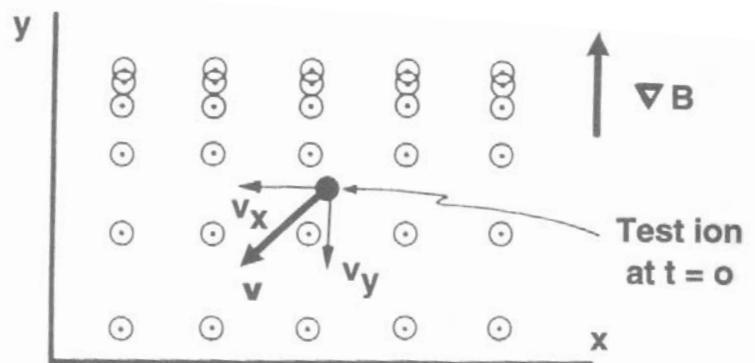
$$\bar{\mathbf{v}}_{gc,\perp} = \frac{\bar{\mathbf{F}} \times \mathbf{B}}{qB^2} = \mathbf{v}_{DF}$$



Individual Charge Trajectories

- Spatially varying magnetic field

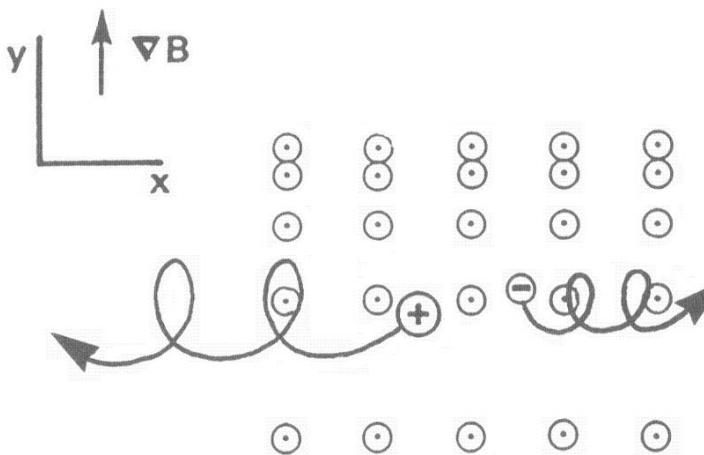
$$m \frac{d\mathbf{v}}{dt} = q[\mathbf{v} \times \mathbf{B}(r)]$$



$$\bar{\mathbf{F}} \approx -|q| \left(\frac{dB}{dy} \right) \left(\frac{v_{\perp}^2}{2\omega_c} \right) \mathbf{j} = -|q| \frac{v_{\perp}^2}{2\omega_c} \nabla B$$

$$= -\frac{|q|}{2} v_{\perp} r_L \nabla B = -\frac{mv_{\perp}^2/2}{B} \nabla B$$

$$\nabla B = \frac{\partial B_z}{\gamma} \mathbf{j}, \quad v_{\parallel} = 0$$



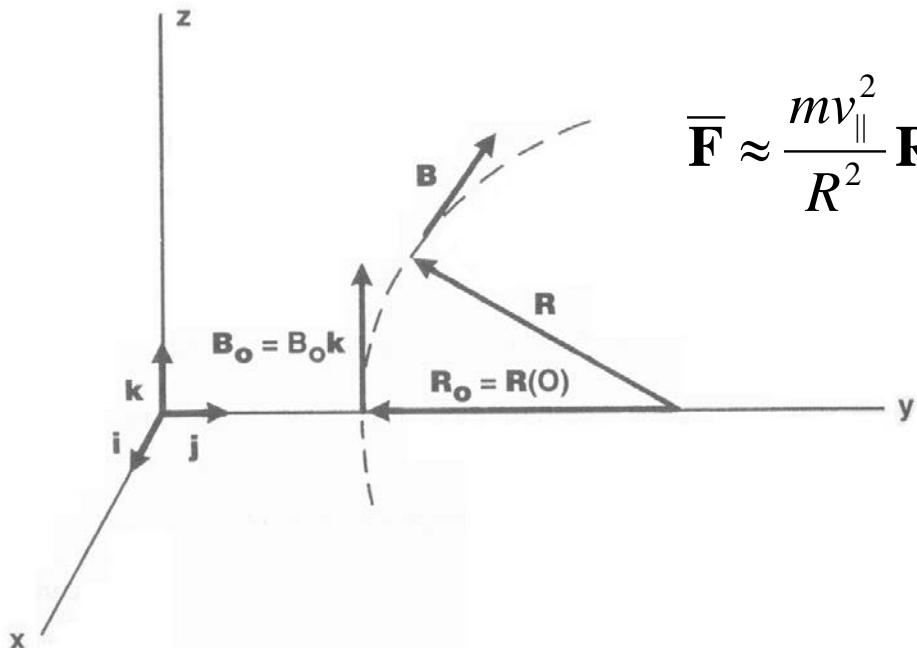
$$\mathbf{v}_{D,\nabla B} = \frac{\bar{\mathbf{F}} \times \mathbf{B}}{qB^2}$$

$$= \pm \frac{v_{\perp}^2}{2\omega_c} \frac{\mathbf{B} \times \nabla B}{B^2} = \pm \frac{1}{2} v_{\perp} r_L \frac{\mathbf{B} \times \nabla B}{B^2}$$

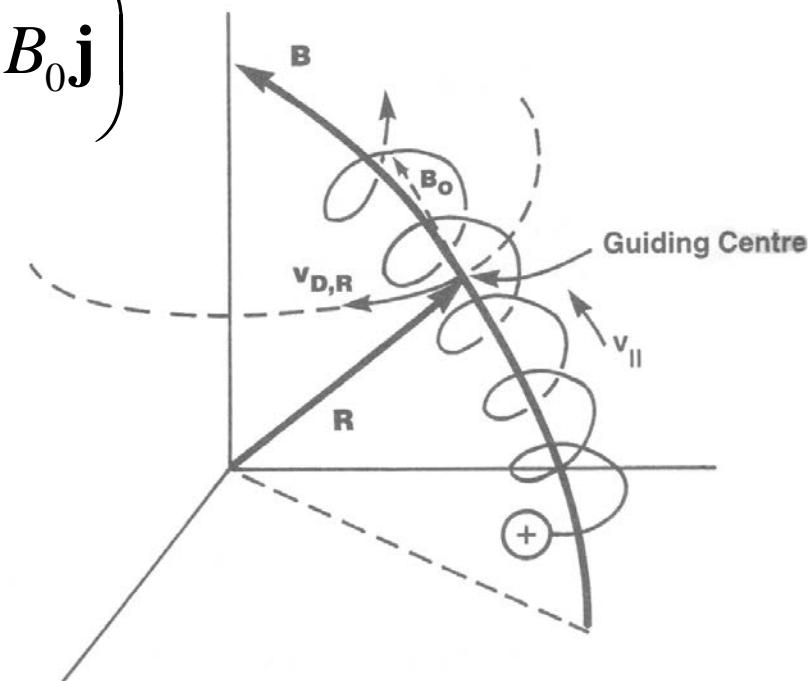
Individual Charge Trajectories

- Curvature drift

$$m \frac{d\mathbf{v}}{dt} = q(\mathbf{v} \times \mathbf{B}) = q(\mathbf{v} \times B_0 \mathbf{k}) + q\left(\mathbf{v} \times \frac{z}{R} B_0 \mathbf{j}\right)$$



$$\bar{\mathbf{F}} \approx \frac{mv_{||}^2}{R^2} \mathbf{R}_0$$



$$\mathbf{v}_{D,R} = \bar{\mathbf{v}}_{gc,x} = \frac{mv_{||}^2}{qB_0^2} \frac{\mathbf{R}_0 \times \mathbf{B}_0}{R^2}$$

Individual Charge Trajectories

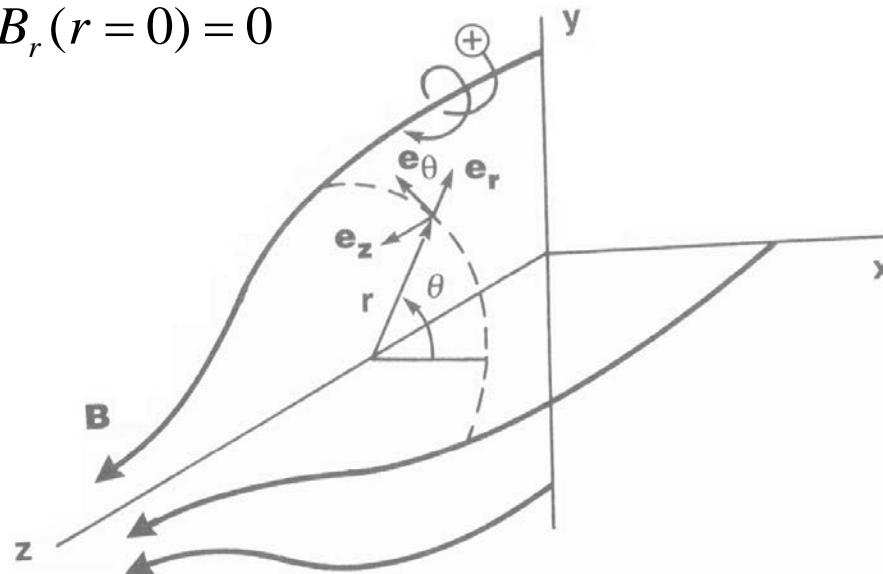
- Axial field variation

$$m \frac{d\mathbf{v}}{dt} = q(\mathbf{v} \times B_z \mathbf{e}_z) + q(\mathbf{v} \times B_r \mathbf{e}_r)$$

$$\nabla \cdot \mathbf{B} = 0 \rightarrow B_r \approx -\frac{r}{2} \frac{\partial B_z}{\partial z} \quad B_\theta = 0, \quad B_r(r=0) = 0$$

$$\mathbf{F}_\parallel = -\frac{1}{2} \frac{mv_\perp^2}{B} \nabla_\parallel B = -\mu \nabla_\parallel B$$

μ : magnetic moment of the gyrating particle



Individual Charge Trajectories

- Invariant of motion

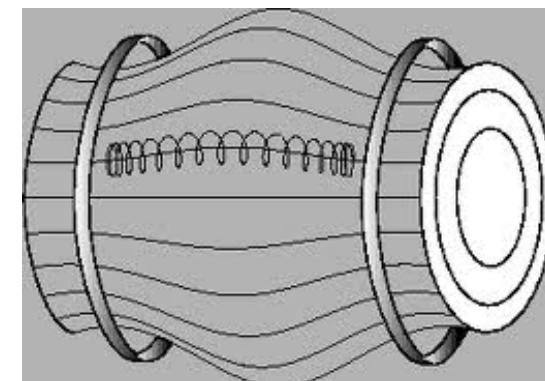
$$\frac{d}{dt} E_0 = \frac{d}{dt} \left(\frac{1}{2} m v_{\perp}^2 + \frac{1}{2} m v_{\parallel}^2 \right) = 0 \quad \mu = \frac{m v_{\perp}^2 / 2}{B} \quad m \frac{dv_{\parallel}}{dt} = -\frac{\mu}{v_{\parallel}} \frac{dB}{dt} \quad m v_{\parallel} \frac{dv_{\parallel}}{dt} = -\mu \frac{dB}{dt}$$

$$\mathbf{F}_{\parallel} = m \frac{d\mathbf{v}_{\parallel}}{dt} = -\mu \nabla_{\parallel} \mathbf{B} = -\mu \frac{\partial \mathbf{B}}{\partial s} = -\mu \frac{\partial \mathbf{B}}{\partial s} \cdot \frac{ds}{dt} \cdot \frac{dt}{ds} = -\mu \frac{\partial \mathbf{B}}{\partial s} \cdot \frac{ds}{dt} \cdot \frac{1}{v_{\parallel}} = -\frac{\mu}{v_{\parallel}} \frac{d\mathbf{B}}{dt} \rightarrow \frac{d}{dt} \left(\frac{1}{2} m v_{\parallel}^2 \right) = -\mu \frac{dB}{dt}$$

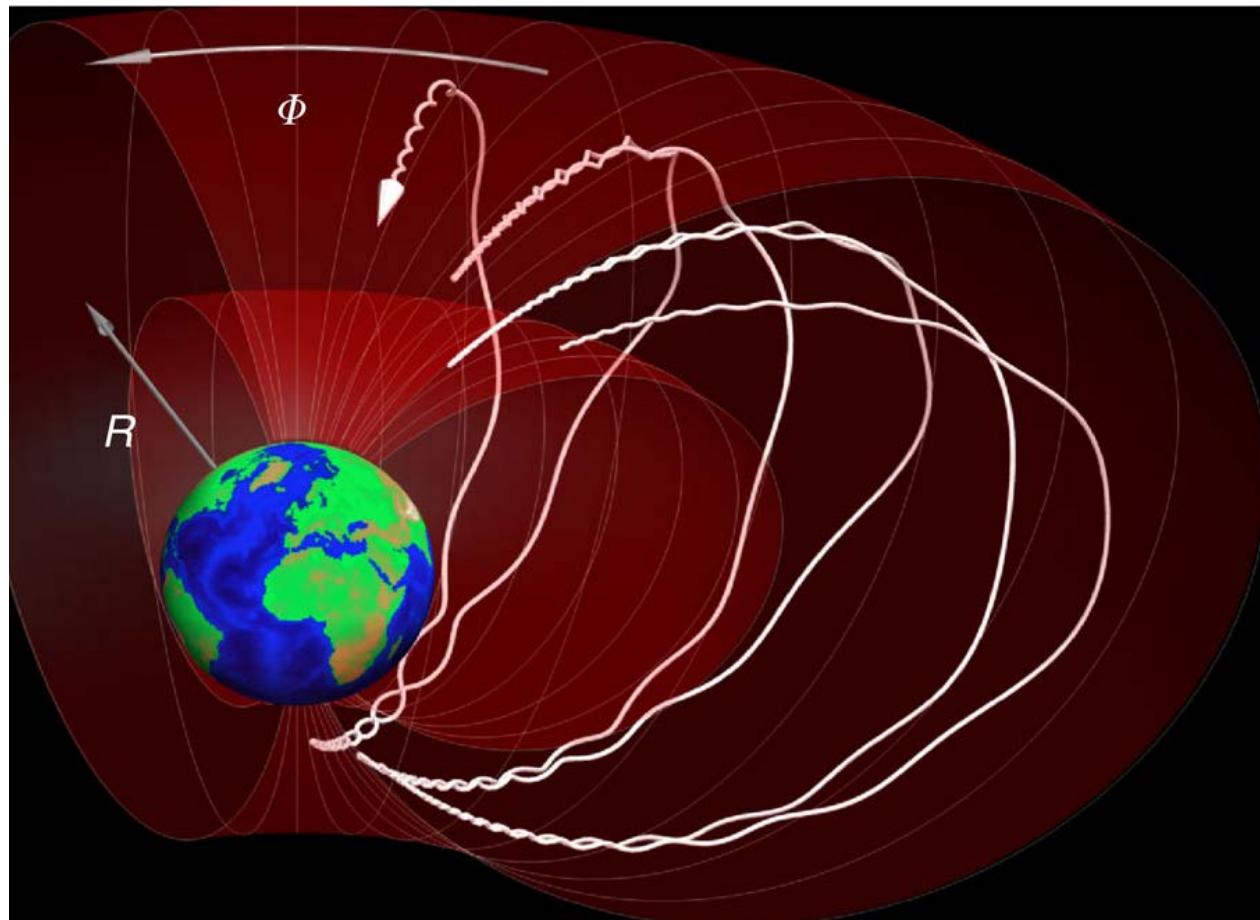
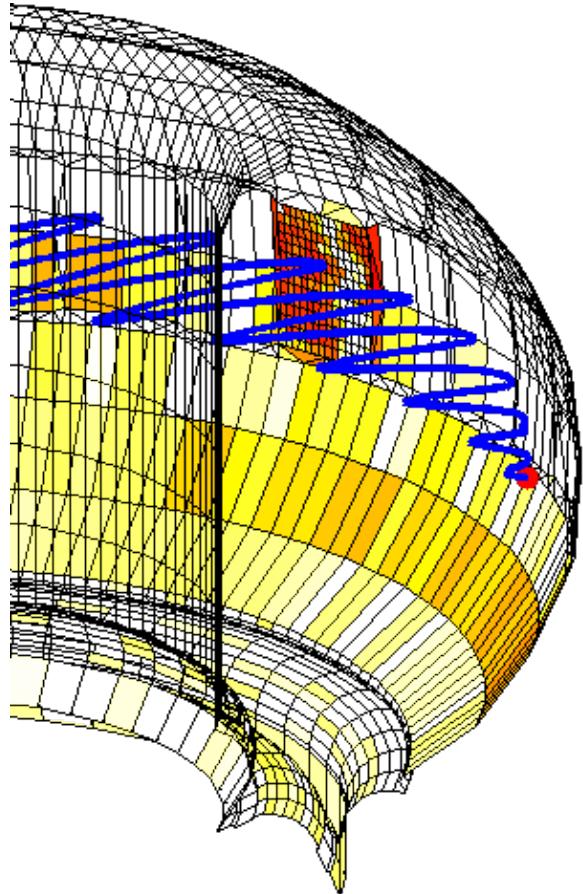
$$\rightarrow \frac{d}{dt} \left(\frac{1}{2} m v_{\perp}^2 + \frac{1}{2} m v_{\parallel}^2 \right) = \frac{d}{dt} (\mu B) + \left(-\mu \frac{dB}{dt} \right) = 0$$

$$\frac{d}{dt} (\mu B) = \mu \frac{dB}{dt} + B \frac{d\mu}{dt}$$

$$\rightarrow \frac{d\mu}{dt} = 0 : \text{adiabatic invariant}$$



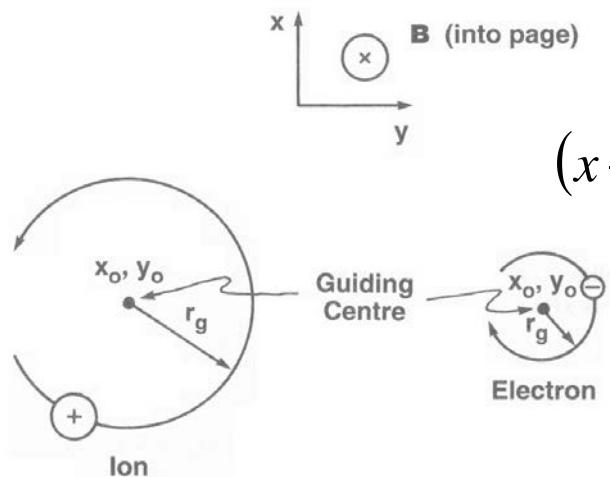
Individual Charge Trajectories



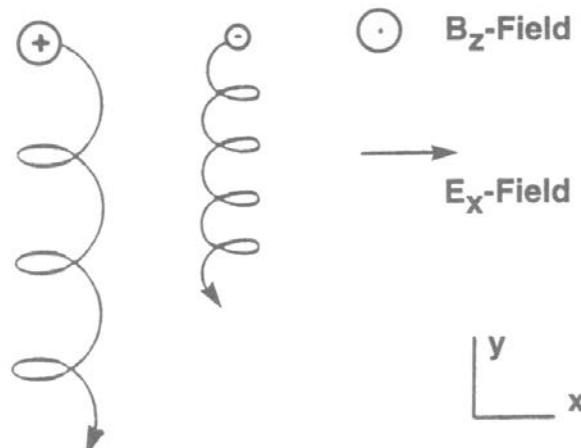
<http://www.physics.ucla.edu/icnsp/Html/spong/spong.htm>

J. P. Graves et al, *Nature Communications* **3** 624 (2012)

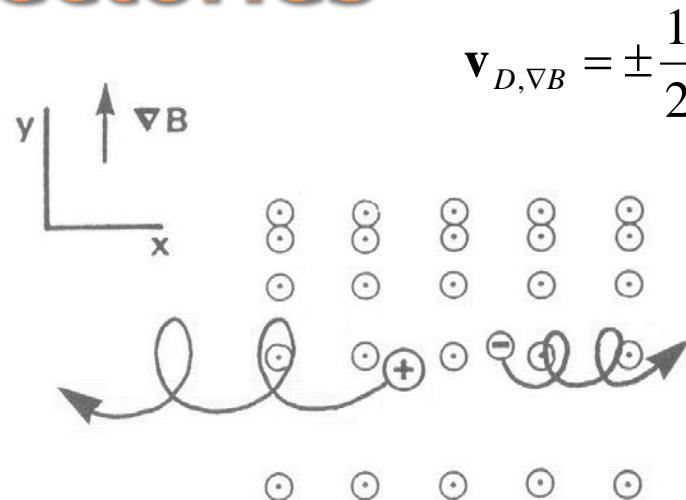
Individual Charge Trajectories



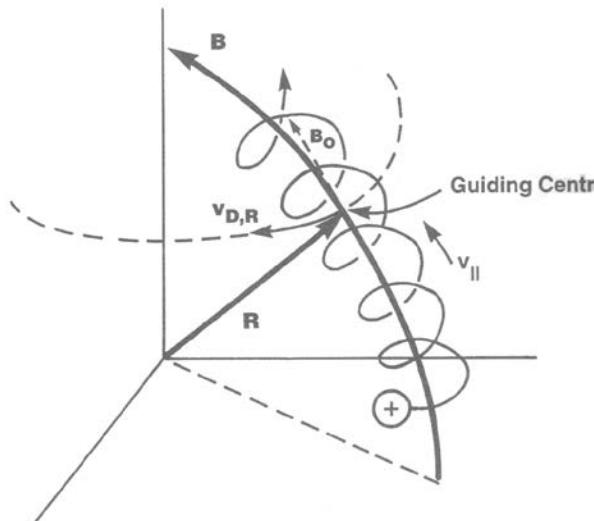
$$(x - x_0)^2 + (y - y_0)^2 = r_L^2$$



$$\mathbf{v}_{D,E} = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$$



$$\mathbf{v}_{D,\nabla B} = \pm \frac{1}{2} v_{\perp} r_L \frac{\mathbf{B} \times \nabla B}{B^2}$$



$$\mathbf{v}_{D,R} = \frac{mv_{||}^2}{qB_0^2} \frac{\mathbf{R}_0 \times \mathbf{B}_0}{R^2}$$

Kinetic Approach of Plasmas

- Boltzmann equation

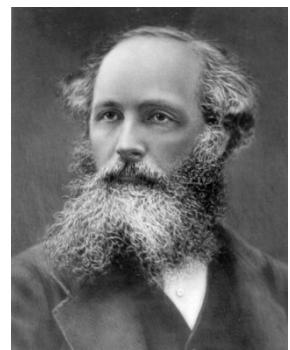
$$\frac{df_\alpha}{dt} = \frac{\partial f_\alpha}{\partial t} + \vec{v} \cdot \nabla f_\alpha + \vec{a} \cdot \nabla_{\vec{v}} f_\alpha = \frac{\partial f_\alpha}{\partial t} + \vec{v} \cdot \nabla f_\alpha + \frac{q_\alpha}{m_\alpha} (\vec{E} + \vec{v} \times \vec{B}) \cdot \nabla_{\vec{v}} f_\alpha = \left(\frac{\partial f_\alpha}{\partial t} \right)_c$$



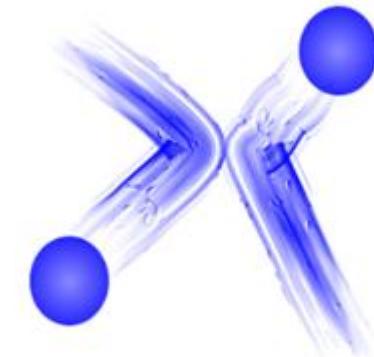
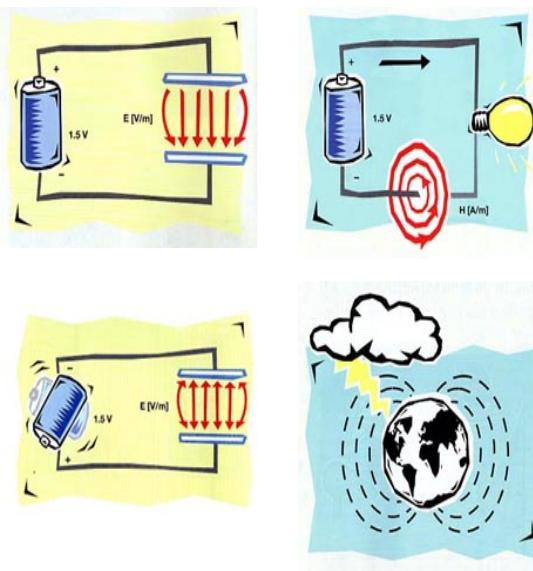
Ludwig Boltzmann
(1844-1906)



WIKIPEDIA



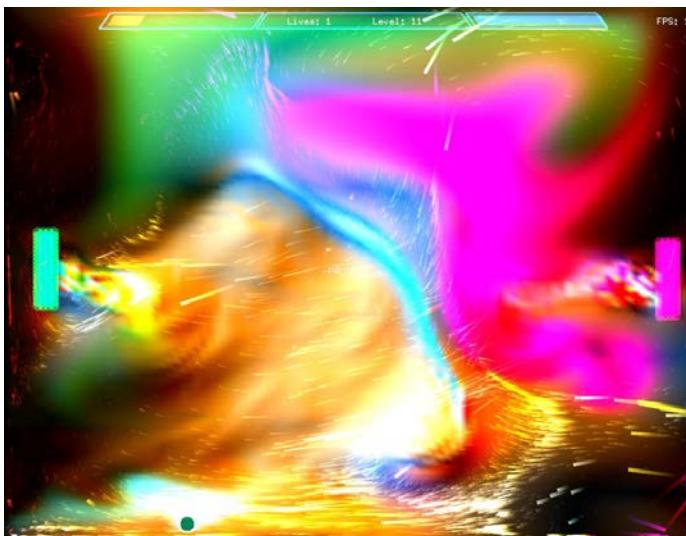
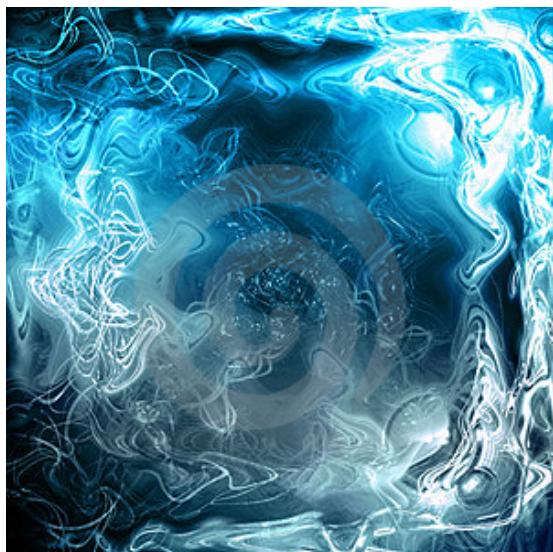
James Clerk Maxwell
(1831-1879)



Fluid Approach of Plasmas

- **Plasmas as fluids**

- The single particle approach gets to be complicated.
- A more statistical approach can be used because we cannot follow each particle separately.
- Introduce the concept of an **electrically charged current-carrying fluid**.



<http://www.tower.com/music-label/gulan-music-studio>
Album art of music OST "Plasma Pong"

Plasmas as Fluids

- Two fluid equations

$$\int Q_i \left[\frac{df_\alpha}{dt} - \left(\frac{\partial f_\alpha}{\partial t} \right)_c \right] d\vec{v} = 0$$

$$Q_1 = 1$$

mass

$$Q_2 = m_\alpha \vec{v}$$

momentum

$$Q_3 = m_\alpha v^2 / 2$$

energy

$$\frac{df_\alpha}{dt} = \frac{\partial f_\alpha}{\partial t} + \vec{v} \cdot \nabla f_\alpha + \frac{q_\alpha}{m_\alpha} (\vec{E} + \vec{v} \times \vec{B}) \cdot \nabla_{\vec{v}} f_\alpha = \left(\frac{\partial f_\alpha}{\partial t} \right)_c$$

$$\left(\frac{\partial f_\alpha}{\partial t} \right)_c = \sum_\beta C_{\alpha\beta}$$

$$\left(\frac{dn_\alpha}{dt} \right)_\alpha + n_\alpha \nabla \cdot \vec{u}_\alpha = 0$$

$$\vec{P}_\alpha \equiv n_\alpha m_\alpha \langle w^2 \rangle$$

$$\vec{v} = \vec{u}_\alpha(\vec{r}, t) + \vec{w}, \quad \langle \vec{w} \rangle = 0$$

$$n_\alpha m_\alpha \left(\frac{d\vec{u}_\alpha}{dt} \right)_\alpha = q_\alpha n_\alpha (\vec{E} + \vec{u}_\alpha \times \vec{B}) - \nabla \cdot \vec{P}_\alpha + \vec{R}_\alpha$$

$$\vec{R}_\alpha \equiv \int m_\alpha \vec{w} C_{\alpha\beta} d\vec{w}$$

$$h_\alpha \equiv \frac{1}{2} n_\alpha m_\alpha \langle w^2 \vec{w} \rangle$$

$$\frac{3}{2} n_\alpha \left(\frac{dT_\alpha}{dt} \right)_\alpha = - \vec{P}_\alpha : \nabla \vec{u}_\alpha - \nabla \cdot \vec{h}_\alpha + Q_\alpha$$

$$Q_\alpha \equiv \int \frac{1}{2} m_\alpha w_\alpha^2 C_{\alpha\beta} d\vec{w}$$

Plasmas as Fluids

- **Single-fluid magnetohydrodynamics (MHDs)**

A single-fluid model of a fully ionised plasma, in which the plasma is treated as a single hydrodynamic fluid acted upon by electric and magnetic forces.

- **The magnetohydrodynamic (MHD) equation**

$$\rho = n_i M + n_e m \approx n(M + m) \approx nM \quad \text{mass density}$$

Hydrogen plasma,
charge neutrality
assumed

$$\sigma = (n_i - n_e)e \quad \text{charge density}$$

$$\vec{v} = (n_i M \vec{u}_i + n_e m \vec{u}_e) / \rho \approx (M \vec{u}_i + m \vec{u}_e) / (M + m) \approx \vec{u}_i + (m/M) \vec{u}_e \quad \text{mass velocity}$$

$$\vec{J} = e(n_i \vec{u}_i - n_e \vec{u}_e) \approx ne(\vec{u}_i - \vec{u}_e) \quad \text{electron inertia neglected:}$$

electrons have an infinitely fast response time because of their small mass

$$\vec{u}_i \approx \vec{v} + \frac{m}{M} \frac{\vec{J}}{ne}, \quad \vec{u}_e \approx \vec{v} - \frac{\vec{J}}{ne}$$

Plasmas as Fluids

F. F. Chen, Ch. 5.7

- Ideal MHD model

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \vec{v} = 0$$

Mass continuity equation
(continuously flowing fluid)

$$\rho \frac{d\vec{v}}{dt} = \vec{J} \times \vec{B} - \nabla p$$

Single-fluid equation of motion

$$\frac{d}{dt} \left(\frac{p}{\rho^\gamma} \right) = 0$$

Energy equation (equation of state):
adiabatic evolution

$$\vec{E} + \vec{v} \times \vec{B} = 0$$

Ohm's law: perfect conductor \rightarrow "ideal" MHD

$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

Maxwell equations

$$\nabla \times \vec{B} = \mu_0 \vec{J}$$

$\epsilon_0 \rightarrow 0$ assumed
(Full \rightarrow low-frequency Maxwell's equations)
Displacement current, net charge neglected

$$\nabla \cdot \vec{B} = 0$$

Plasmas as Fluids

- Magnetohydrodynamics (MHD): 1946 by H. Alfvén

STOCKHOLMS OBSERVATORIUMS ANNALER
(ASTRONOMiska IAKTTAGELSER OCH UNDERSÖKNINGAR Å STOCKHOLMS OBSERVATORIUM)
BAND 14. N:o 9.

ON THE COSMOGONY OF THE SOLAR SYSTEM

III

BY

HANNES ALFVÉN



Hannes Alfvén
(1908-1995)

"Nobel prize in
Physics (1970)"

Plasmas as Fluids

- Magnetohydrodynamics (MHD): 1946 by H. Alfvén

STOCKHOLMS OBSERVATORIUMS ANNALER. BAND 14. N:o 9.

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The laws found can be applied to the planetary as well as to the satellite systems (§§ 7—9) and also — as is pointed out in § 10 — to the non-solar planets newly discovered.

At last some remarks are made about the transfer of momentum from the Sun to the planets, which is fundamental to the theory (§ 11). The importance of the magnetohydrodynamic waves in this respect is pointed out. It is possible that the Sun's general magnetic field was much stronger at the time of formation of the solar system than it is now. This must be assumed in order to explain the momentum transfer quantitatively.

Kungl. Tekniska Högskolan, Stockholm, November 1945.



Plasmas as Fluids

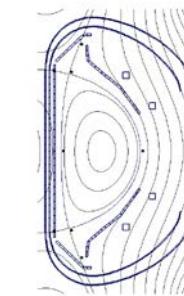
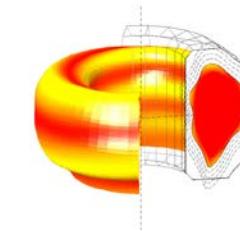
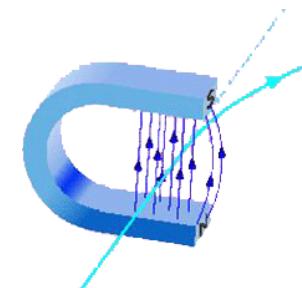
- Ideal MHD

- Single-fluid model
- Ideal:
 - Perfect conductor with zero resistivity

- MHD:
 - Magnetohydrodynamic (magnetic fluid dynamic)

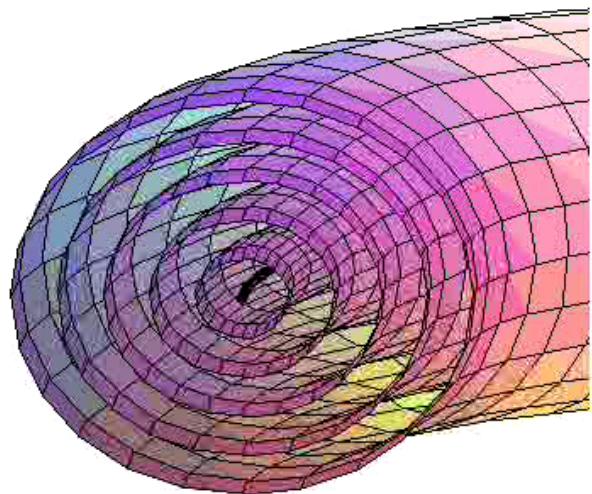
- Assumptions:
 - Low-frequency, long-wavelength
 - collision-dominated plasma

- Applications:
 - Equilibrium and stability in fusion plasmas

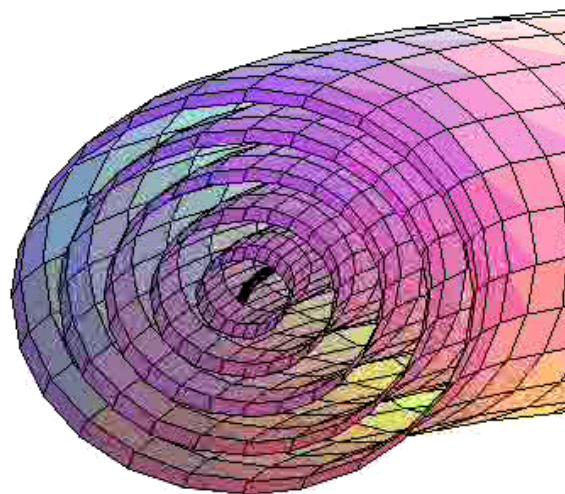


What is Ideal MHD?

- Ideal MHD: $\eta = 0$

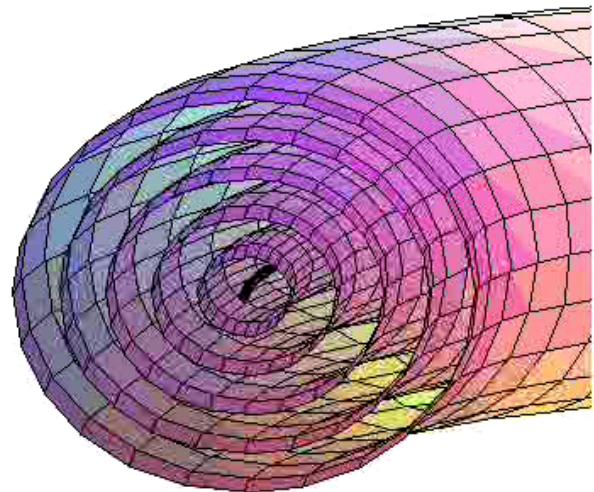


- Resistive MHD: $\eta \neq 0$

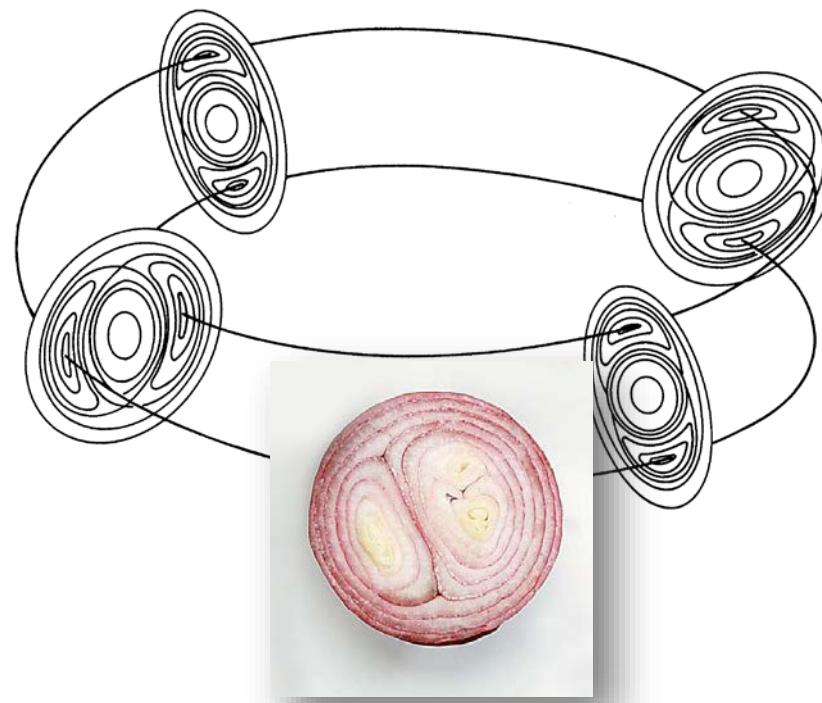


What is Ideal MHD?

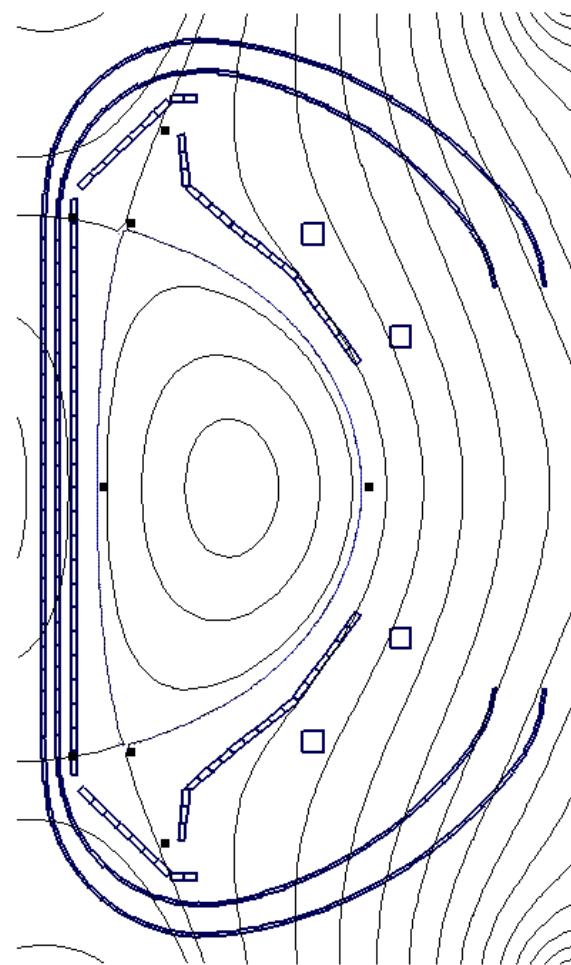
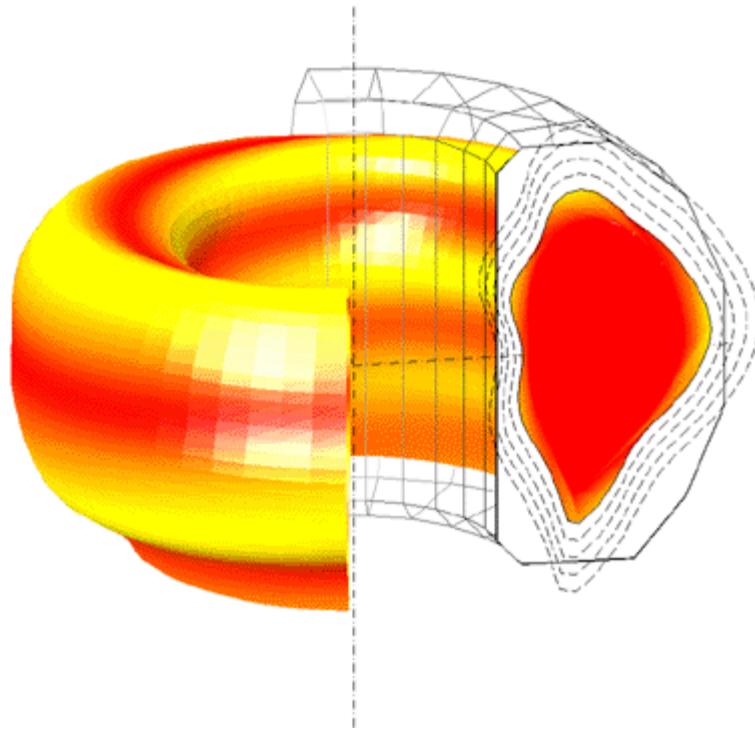
- Ideal MHD: $\eta = 0$



- Resistive MHD: $\eta \neq 0$



Applications



Plasma Eqaulilibrium and Stability