

Fusion Plasma Theory II. 2019

Week 2

Ch. 16. Waves in an unmagnetized plasma

II.-1

16.1. Langmuir Waves

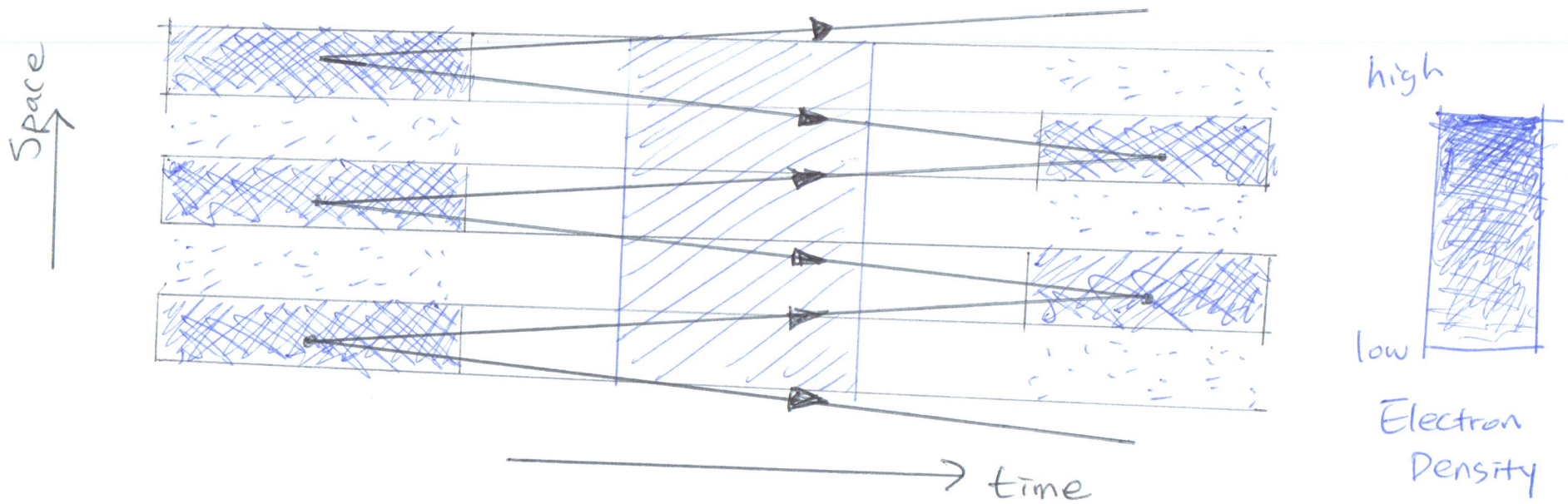


Fig 16.1.

- * Initial displacements of electrons from equilibrium position
 - charge separation → electric field generation
 - pulling electrons back → overshoot → oscillations,

* This oscillation happens in very short time scale

\Rightarrow ions with heavy mass cannot respond,
and can be considered as stationary.

* Electron Dynamics :

$$- m n_e [\dot{\vec{u}}_e + (\vec{u}_e \cdot \vec{\nabla}) \vec{u}_e] = -e n_e \vec{E} - \vec{\nabla} p_e \quad (16.1)$$

$$- \dot{n}_e + \vec{\nabla} \cdot (n_e \vec{u}_e) = 0 \quad (16.2)$$

$$- \epsilon_0 \vec{\nabla} \cdot \vec{E} = e (n_i - n_e) \quad (16.3)$$

* Consider electrons moving in x direction and
waves propagating " " .

\Rightarrow perturbed quantities vary as "exp [i (kx - \omega t)]"

$$\boxed{\frac{\partial}{\partial x} \rightarrow i k} \quad \text{and} \quad \boxed{\frac{\partial}{\partial t} = -i \omega}$$

* Linearize Eqs. (16.1)-(16.3) i.e., keep only the 1st order quantities:

In addition, relate p_1 to n_1 by using the equation of state for 1-dimensional compression faster than the thermal conduction, i.e., $p \propto n^\gamma$, where $\gamma = \frac{2+N}{N} = \frac{2+1}{1} = 3$.

$$\Rightarrow \left[\begin{array}{l} i\omega m n_0 u_1 = e n_0 E_1 + 3 i k T n_1 \end{array} \right. \quad (16.6)$$

$$\left[\begin{array}{l} -i\omega n_1 + i k n_0 u_1 = 0. \end{array} \right. \quad (16.7)$$

$$\left[\begin{array}{l} i k \epsilon_0 E_1 = -e n_1 \end{array} \right. \quad (16.8)$$

3 equations for 3 perturbed quantities, u_1 , n_1 and E_1 .

* $[\text{Determinant}] = 0$
of 3×3 matrix.

\Rightarrow Dispersion Relation
(between " ω " and " k ")

*

$$\omega^2 = \omega_{pe}^2 + 3k^2 v_{Te}^2$$

where

$$\omega_{pe}^2 \equiv n_e e^2 / \epsilon_0 m_e$$

electron plasma frequency

$$v_{Te}^2 = (T_e / m_e)$$

electron thermal velocity.

Bohm-Gross dispersion relation [PR 1949]

* Fig 16.2 illustrates the behavior at different k regimes.

⊗ At long wavelength (low k) ; $\omega^2 \approx \omega_{pe}^2$

~ non-propagating "plasma oscillation"
(Langmuir)

⊗ At short wavelength (high k) ;

$$\omega^2 \approx 3k^2 v_{Te}^2$$

~ Propagating "electron sound wave."

16.2. Ion Sound Waves

II-5.

* Take a look at another electrostatic longitudinal wave

($\vec{k} \parallel \vec{E}_1$) which oscillates in longer time scale

so that ions dynamics plays an essential role.

→ Electrons can move fast enough to establish nearly exact force balance (between pressure gradient and electric field), i.e., a Boltzmann response.

* Follow a similar derivation as the one for Langmuir wave.

- Ion Fluid momentum eqn. (16.16)

- Poisson eqn. (16.19)

- Ion continuity eqn (16.21)

* A crucial difference (from the Langmuir wave where $n_{i1} = 0$)

is that

$$n_{e1} = n_e - n_{e0} = n_{e0} (\exp(e\phi_1/T_e) - 1) \approx n_{e0} e\phi_1/T_e \quad (16.18)$$

i.e. Boltzmann response (or adiabatic response),

* Linearization \rightarrow Solve ~~the~~ coupled eqns for n_{i1} , ϕ_1 , and $U_{i1} \rightarrow$ to get:

$$\left(\omega/k\right)^2 = \frac{T_e/M}{1 + k^2 \lambda_D^2} + \gamma_i T_i/M \quad (16.23)$$

$$\lambda_D^2 \equiv \epsilon_0 T_e / n_e e^2 = v_{Te}^2 / \omega_p^2$$

Debye length.

* In the long wavelength limit (~~low~~ low k);

$$\star \quad \left(\omega/k \right)^2 \approx \frac{T_e + \gamma_i T_i}{M} \quad \text{Sound wave}$$

— Both electron and ion pressure contribute,
but only ions provide 'mass'.

— In collisionless plasma, Eq. \star is valid for $T_e \gg T_i$

because one needs to consider wave-particle
resonant ~~ex~~ interaction (Landau damping) for

$\omega/k \sim v_{Ti}$ using kinetic theory.

* In the short wavelength limit (high k);

$$\omega^2 \approx \frac{T_e/M}{\lambda_D^2} \approx \left(\frac{m}{M} \right) \omega_p^2 \equiv \Omega_p^2; \quad \text{ion plasma frequency.}$$

non-propagating wave.

16.3. High-Frequency Electromagnetic Waves in an Unmagnetized Plasma.

II-8.

* Now, we consider magnetic field perturbation \vec{B}_1 in addition to \vec{E}_1 . Then, we ~~are~~ need to consider Ampère's law and Faraday's law. With $\vec{\nabla} \Rightarrow i\vec{k}$,

$$- i\vec{k} \times \vec{B}_1 = \mu_0 \vec{J}_1 - i\omega \vec{E}_1 / c^2 \quad (16.24)$$

$$- i\vec{k} \times \vec{E}_1 = i\omega \vec{B}_1 \quad (16.25)$$

⊗ Taking $\vec{k} \times$ Eq. (16.25) and expanding the triple vector product,

$$\boxed{k^2 \vec{E}_1 - \vec{k} (\vec{k} \cdot \vec{E}_1) = \left(\frac{\omega}{c}\right)^2 (\vec{E}_1 + i \vec{J}_1 / \epsilon_0 \omega)} \quad (16.27)$$

* For longitudinal waves with $\vec{k} \parallel \vec{E}_1$, LHS = 0 and we have an electrostatic wave in which Poisson eqn plays an important role.

* We consider here a transverse wave with $\vec{k} \perp \vec{E}_1$

(i.e., $\vec{k} \cdot \vec{E}_1 = 0$). \Rightarrow no need to consider Poisson eqn or density perturbation or continuity eqn.

* $\boxed{\vec{J}_1 = -n_0 e \vec{u}_1} \quad (16.28)$

\swarrow
where

(assume stationary ions
for high frequency waves
as before).

$\boxed{-i\omega m \vec{u}_1 = -e \vec{E}_1} \quad (16.29)$

\Rightarrow Eq. (16.27) becomes,

$$(c^2 k^2 - \omega^2) \vec{E}_1 = i\omega \vec{J}_1 / \epsilon_0 = -\left(\frac{n_0 e^2}{m \epsilon_0}\right) \vec{E}_1 \quad (16.31)$$

$$\omega^2 = c^2 k^2 + \omega_p^2$$

(16.32)

Dispersion Relation for "EM wave" in unmagnetized plasma.

(*) Note that:

$$v_p \equiv \frac{\omega}{k} = c \left(1 + \frac{\omega_p^2}{c^2 k^2} \right)^{1/2} > c.$$

while

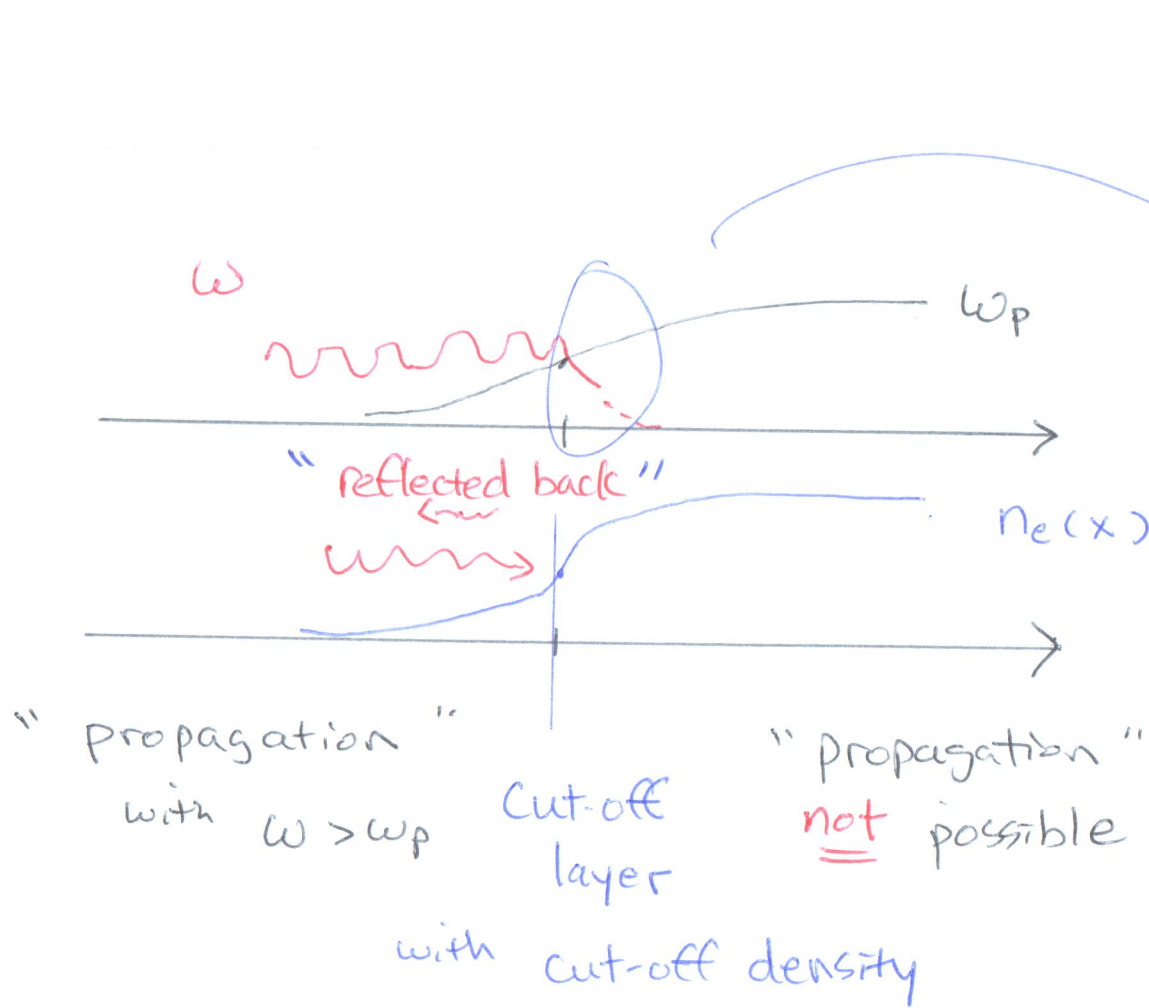
$$v_g \equiv \frac{\partial \omega}{\partial k} = \frac{c^2}{v_p} < c.$$

This is **NOT** a violation of the special relativity because the information and energy propagate with v_g **not** v_p .

- low k ; \leadsto constant freq, non-propagating ~~oscillation~~
- hi k ; \leadsto EM waves in vacuum propagating with "c".

(*) Homework Problem 16.2 on page 266.

- (*) EM waves cannot propagate in a plasma with
 $\omega_p > \omega$.



$$n_c = m_e \epsilon_0 \omega^2 / e^2.$$

For $\omega_p > \omega$,
 D.R. admits "k"
 which is imaginary,
 i.e.,
 $k = (\omega^2 - \omega_p^2)^{1/2} / c$
 $= \pm i (\omega_p^2 - \omega^2)^{1/2} / c$

c(16.37)

$$\exp(i k x) = \exp\left(-x \frac{(\omega_p^2 - \omega^2)^{1/2}}{c}\right)$$

evanescent solution

• Collisionless skin depth:
 $\equiv "c/\omega_p"$