

Fusion Plasma Theory II. 2019

Week 3

Ch. 17. High-Frequency Waves in Magnetized Plasmas

- ⊗ In the presence of a strong equilibrium magnetic field, \vec{B}_0 ,
the medium is anisotropic.

17.1. High-Frequency EM waves propagating \perp to \vec{B}_0

⊗ $\vec{k} \perp \vec{B}_0$:

- ①. O-waves: $\vec{E}_1 \parallel \vec{B}_0 \rightarrow \vec{B}_0$ plays no role in wave dynamics.
"ordinary"

\Rightarrow we can apply the results for unmagnetized plasma with $\vec{u}_1 \parallel \vec{E}_1$,

$$\vec{u}_1 \times \vec{B}_0 = 0 \quad \text{and} \quad \text{O-mode (O-wave) never notices } \vec{B}_0.$$

\Rightarrow Recall that EM waves in unmagnetized plasma is "transverse".
" $\vec{E}_1 \perp \vec{k}$ ", with \perp propagation ($\vec{k} \perp \vec{B}_0$) ~~$\rightarrow \vec{E}_1 \perp \vec{k}$~~ .

$$\textcircled{*} \quad \vec{k} \perp \vec{B}_0 :$$

$$\textcircled{2} \text{ X-waves : } \quad \vec{E}_1 \perp \vec{B}_0$$

extra-ordinary ---

- In general, \vec{E}_1 of this X-wave has a component along \vec{k} (\perp to \vec{B}_0) and also a component \perp to both \vec{k} and \vec{B}_0 .

$$\textcircled{*} \text{ For } \vec{B}_0 = B_0 \hat{z} \text{ and } \vec{k} = k \hat{x},$$

$$\vec{E}_1 = E_{1x} \hat{x} + E_{1y} \hat{y}.$$

- For this hi-freq. wave, we will take the ions to be stationary.
- In addition, we will neglect the electron pressure assuming a cold plasma with $T_i \cong T_e \cong 0$.

(*) Recall the wave equation, (16.27):

$$k^2 \vec{E}_1 - \vec{k} (\vec{k} \cdot \vec{E}_1) = \left(\frac{\omega}{c}\right)^2 (\vec{E}_1 + i \vec{j}_1 / \epsilon_0 \omega), \quad (17.4)$$

where $\vec{j}_1 = -n_0 e \vec{u}_1$ carried by electrons.

* \vec{u}_1 can be obtained from the linearized electron fluid eqn of motion including the Lorentz force, $-e(\vec{u}_1 \times \vec{B}_0)$,

$$-i\omega m u_{x1} = -e(E_{x1} + u_{y1} B_0) \quad (17.1)$$

$$-i\omega m u_{y1} = -e(E_{y1} - u_{x1} B_0) \quad (17.2)$$

* This can be solved for u_{x1} and u_{y1} , which will be used in Eq. (17.4)..

$$u_{x1} = \frac{\left(\frac{e}{m}\right) (i\omega E_{x1} + \omega_c E_{y1})}{\omega_c^2 - \omega^2} \quad (17.3)$$

$$u_{y1} = \frac{\left(\frac{e}{m}\right) (i\omega E_{y1} - \omega_c E_{x1})}{\omega_c^2 - \omega^2}, \quad \text{where } \omega_c = \frac{eB_0}{m} \text{ is the electron cyclotron freq.}$$

⊕ From this procedure, we can obtain two coupled eqns

for E_{x1} and E_{y1} .

Determinant of 2×2 matrix = 0
(of coefficients)

\Rightarrow after some arrangements;

\perp propⁿ,
X-mode

$$\frac{c^2 k^2}{\omega^2} = \frac{c^2}{v_p^2} = 1 - \frac{\omega_p^2 (\omega^2 - \omega_p^2)}{\omega^2 (\omega^2 - \omega_h^2)}$$

(17.12)

where $\omega_h^2 \equiv \omega_p^2 + \omega_c^2$, "upper-hybrid" frequency

⊗ Resonance:

- At $\omega = \omega_h$, $k \rightarrow \infty$ ($\lambda \rightarrow 0$), $v_p \rightarrow 0$ and
wave fronts pile up.

- We can also check that $E_{y1}/E_{x1} \rightarrow 0$ at $\omega = \omega_h$,
 $\vec{E}_1 = E_{x1} \hat{x} \parallel \vec{k} = k \hat{x}$ at resonance, and wave is electrostatic at resonance.

(*) Cutoffs ; where $k \rightarrow 0$ ($\lambda \rightarrow \infty$).

III-5.

From Eq. (17.12), numerator on RHS = 0 for $k = 0$,

\Rightarrow

$$\omega = [(\omega_c^2 + 4\omega_p^2)^{1/2} \pm \omega_c] \equiv \begin{cases} \omega_R \\ \omega_L \end{cases} \quad (17.19)$$

(*) Examination of Eq. (17.12); \Rightarrow

Waves can propagate for

but cannot "

$\omega > \omega_R$ and $\omega_h > \omega > \omega_L$,

$\omega_R > \omega > \omega_L$ and $\omega_L > \omega$.

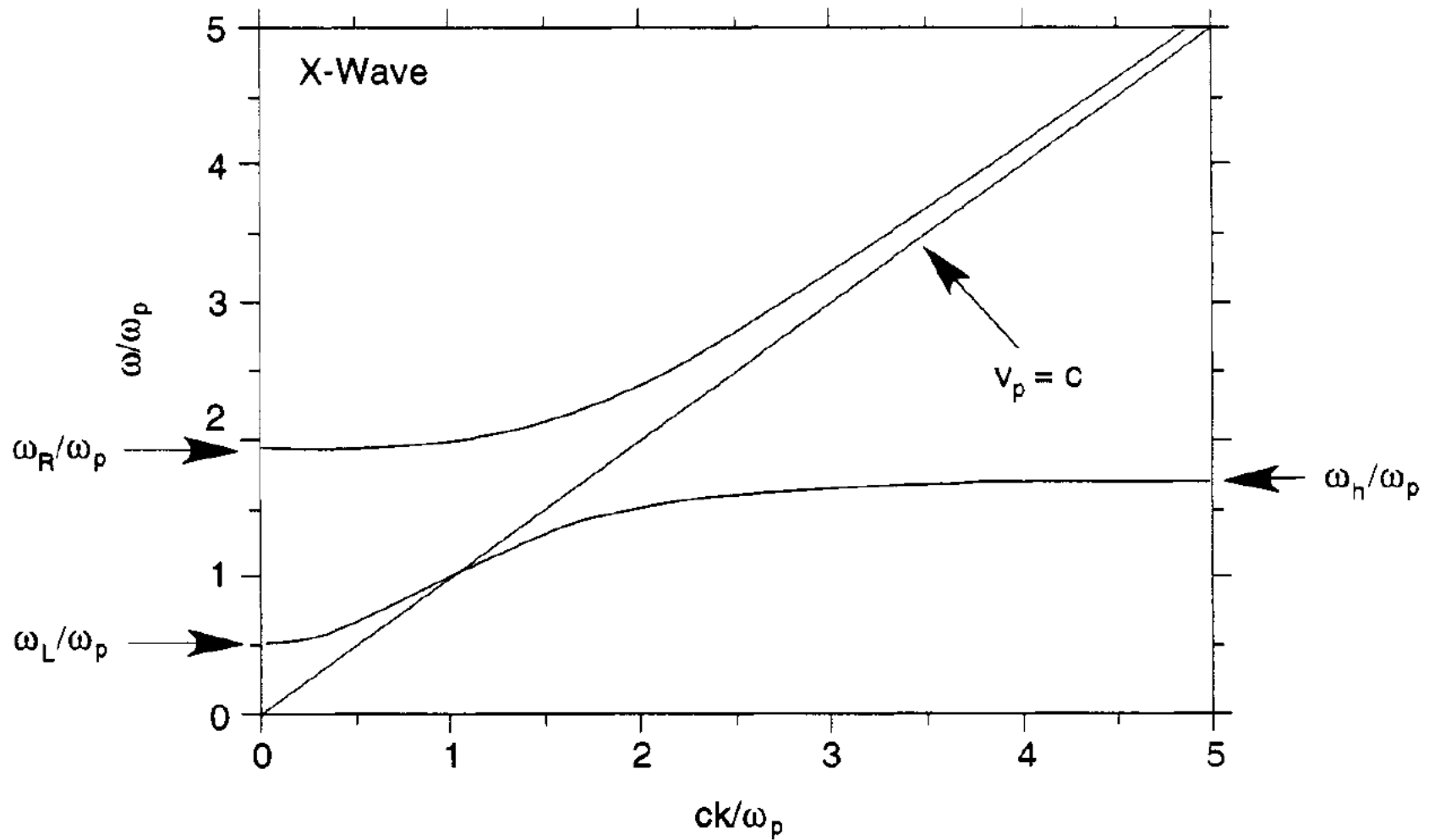


Figure 17.1. Dispersion relation for the extraordinary wave propagating perpendicular to \mathbf{B} in a magnetized plasma, with ω_c^2 chosen to be equal to $2\omega_p^2$.

⊗

Discussion related to Experiments:

- In general, a fixed-frequency wave is driven by a generator.
- It is much easier to arrange a wave propagating up a density gradient, since the RF source is usually located outside of the plasma.
- It is possible to arrange a wave propagating down a \vec{B}_0 gradient (high-field side launch in tokamak).

Recall that $\omega_p^2 \propto n_e$ and $\omega_c \propto |\vec{B}_0|$.

\Rightarrow It's easy to arrange a wave ^{to} propagate and reach ω_R -cutoff.

\Rightarrow To make a wave propagate and reach ω_h -resonance, one should rely on $|\vec{B}_0|$ variation in space.

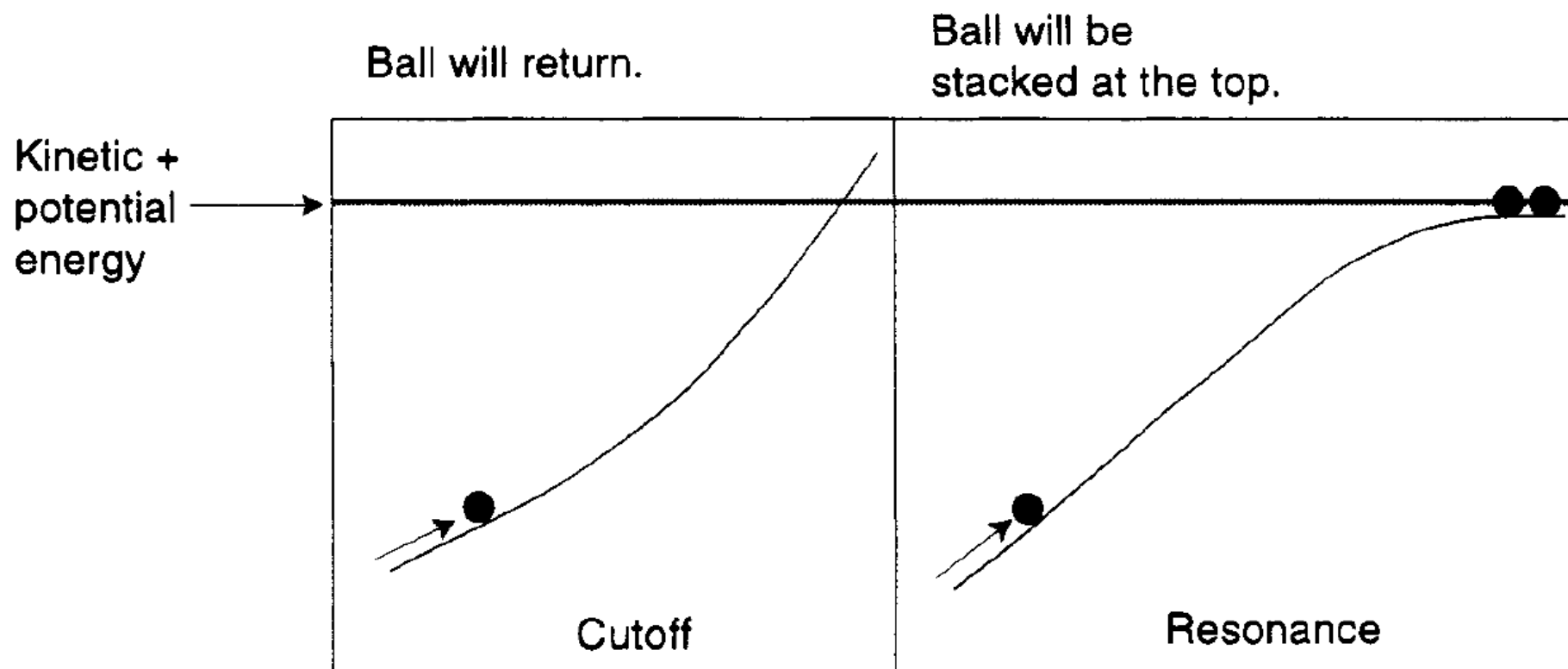


Figure 17.2. Mechanical analog to wave cutoffs and resonances.

- ⊗ Within the context of \perp propagation of EM waves based on linear cold-plasma theory, the wave amplitude must grow steadily at the upper-hybrid resonance layer when we pump energy from the outside.
- ⊗ On the other hand, it can be shown that the wave energy is reflected at the cut-off (accelerating the wave back out of the plasma) \uparrow at $\omega = \omega_p$ in an unmagnetized plasma.
(by using an EM wave example)

Homework : Problem 17.2 on page 277

* 17.2. Hi-Freq. EM waves propagating $\parallel \vec{B}_0$

III-10.

- (*) Consider $\vec{k} \parallel \vec{B}_0$ and high frequency limit such that ions can be considered to be stationary.

Once again, we use the wave eqn.

$$k^2 \vec{E}_1 - \vec{k}(\vec{k} \cdot \vec{E}_1) = \left(\frac{\omega}{c}\right)^2 \left[\vec{E}_1 + i \frac{\vec{j}_1}{\epsilon_0 \omega} \right]. \quad (17.27)$$

- (*) Since a longitudinal mode ($\vec{E}_1 \parallel \vec{k}$) corresponds to the electrostatic Langmuir wave, we consider a new EM wave with $\vec{k} \cdot \vec{E}_1 = 0$ (transverse wave).

- (*) Once again, we can solve the linearized electron fluid eqn of motion, to get ~~the~~ expressions for u_{x1} and u_{y1} in terms of E_{x1} and E_{y1} and substitute to $\vec{j}_1 = -n_e e \vec{U}_1$.

(*) Once again, we get a coupled set of eqns for E_{x1} and E_{y1} , III.-11,

$$\begin{bmatrix} \text{matrix} \end{bmatrix} \begin{bmatrix} E_{x1} \\ E_{y1} \end{bmatrix} = 0. \quad \text{Det} [\text{matrix}] = 0 \quad \text{yields,}$$

$$\tilde{n}^2 \equiv \frac{c^2 k^2}{\omega^2} = 1 - \frac{\omega_p^2}{\omega(\omega \pm \omega_c)} \quad (17.35)$$

\tilde{n} : the index of refraction

- "+" and "-" signs correspond to "L-wave" and "R-wave" respectively.

- Both correspond to "circularly polarized waves."

- E_{x1} and E_{y1} are $\pi/2$ out of phase.

[L wave's \vec{E}_1 vector rotates according to a left-hand rule.
R right-hand rule.

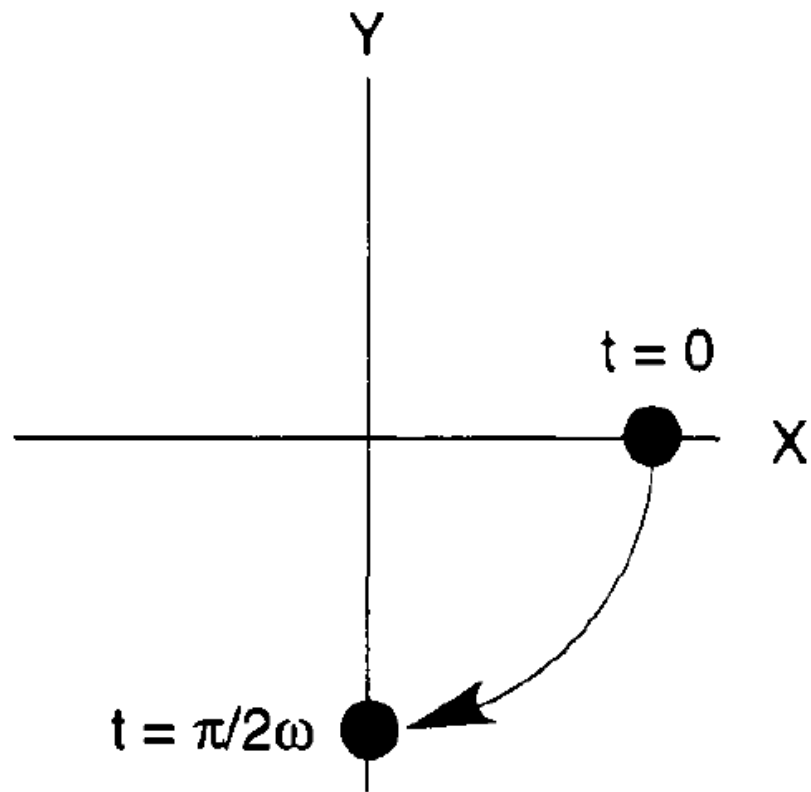


Figure 17.3. Time progression of the **E**-field vector for a left-hand circularly polarized wave with \mathbf{B}_0 along the z direction, out of the page.

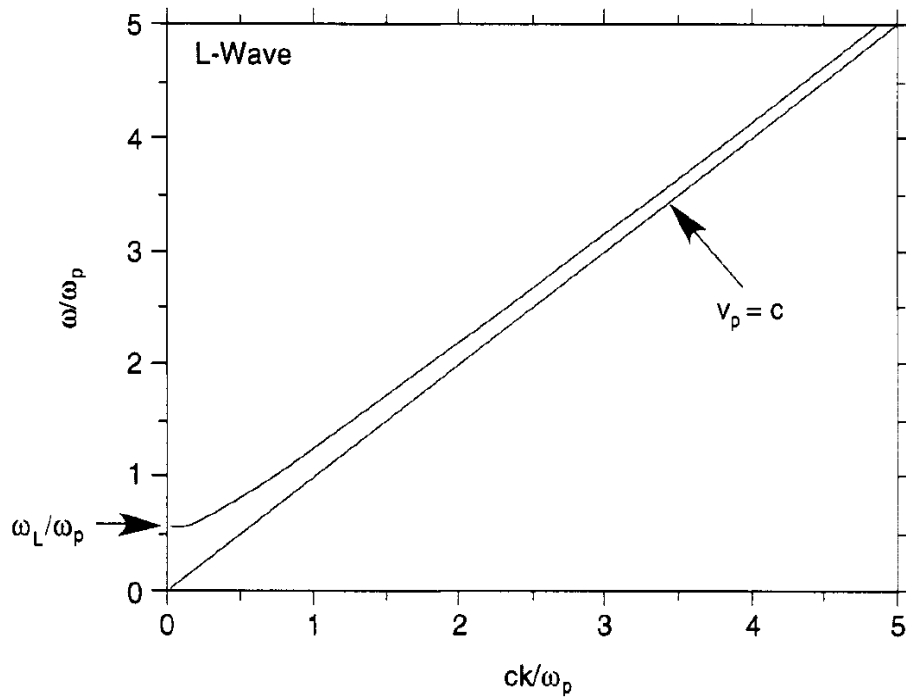


Figure 17.4. Left-hand circularly polarized electromagnetic wave propagating parallel to \mathbf{B}_0 in a magnetized plasma, with ω_c^2 chosen to equal $2\omega_p^2$.

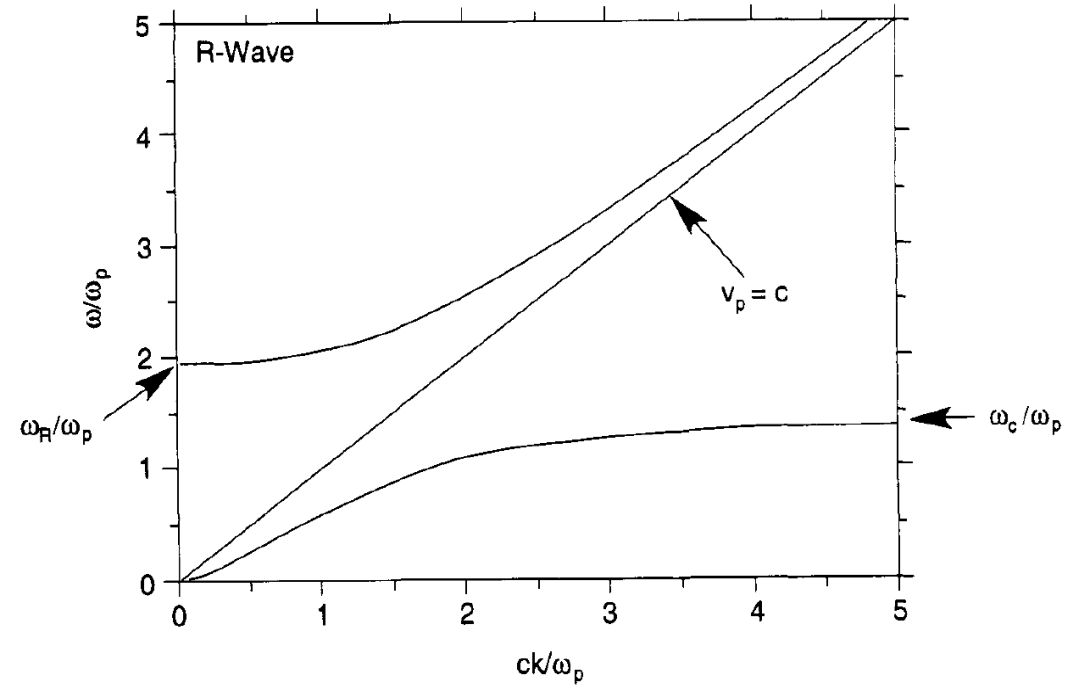


Figure 17.5. Right-hand circularly polarized electromagnetic wave propagating parallel to \mathbf{B}_0 in a magnetized plasma, with ω_c^2 chosen to equal $2\omega_p^2$.

(*) Whistler wave :

(*) R-wave in the region of $\omega < \omega_c$.
(lower frequency range)

(*) One can show that $v_g = \frac{\partial \omega}{\partial k} \nearrow$ as $\omega \nearrow$.

radio
→ White noise generated in a burst ~~at~~ in the ionosphere due to lightning ~~are~~ flashes, and propagating as a whistler wave will travel faster at high frequencies than a low.

⇒ Ground-based receiver will then hear a "whistle" going from hi-freq. to low due to lightning flashes.

⊗ Faraday Rotation:

⊗ For the same freq. " ω ", the upper band of the R-wave ($\omega > \omega_R$) has a higher phase velocity v_p than the corresponding L-wave at the same " ω ".

⇒ When a linearly polarized EM wave propagates \parallel to \vec{B}_0 , the angle of polarization of the wave rotates as it travels.

⊗ This is illustrated on page 283 of G and R.

⊗ Since the resulting difference in the phase depends on ω_c and ω_p , one can determine the magnetic field in plasma if density of plasma is known from other means.

For details, try Problem 17.1 and 17.4.