# Ch. 12 Kinetics of Particles 

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### 12.0 Introduction

Newton's $\left\{\begin{array}{l}1^{\text {st }} \text { law } \rightarrow \\ 3^{\text {rd }} \text { law } \rightarrow\end{array}\right.$ motion of bodies with no acceleration
$2^{\text {nd }}$ law $\rightarrow$ accelerated : magnitude or direction of the velocity changes

Resultant of the forces are not zero, particle will have an acceleration ratio of the resultant force and the acceleration $\rightarrow$ mass

Linear momentum : $\overrightarrow{\boldsymbol{L}}=\boldsymbol{m} \overrightarrow{\boldsymbol{v}}$, alternative form of Newton's $2^{\text {nd }}$ law
International System of Units (SI) vs. US customary units
Rectangular components vs. tangential/normal component
Radial/transverse component
angular momentum $\overrightarrow{\boldsymbol{H}_{0}}=\overrightarrow{\boldsymbol{r}} \times \boldsymbol{m} \overrightarrow{\boldsymbol{v}}$, another Newton's $2^{\text {nd }}$ law
Under central force $\rightarrow$ angular momentum about O is conserved
orbit motion under gravitational attraction

### 12.1A Newton's $2^{\text {nd }}$ Law of Motion

If the resultant force acting on a particle is not zero, the particle will have an acceleration proportional to the magnitude of the resultant and in the direction of this resultant force
Fig. 12.1 (a)~(c)

$$
\begin{gather*}
\frac{F_{1}}{a_{1}}=\frac{F_{2}}{a_{2}}=\frac{F_{3}}{a_{3}}=\cdots=\text { const. } \rightarrow \text { mass } m \\
\vec{F}=m \vec{a} \tag{12.1}
\end{gather*}
$$

$\begin{cases}\overrightarrow{\boldsymbol{F}} & \text { and } \overrightarrow{\boldsymbol{a}} \text { are proportional } \\ \overrightarrow{\boldsymbol{F}} & \text { and } \overrightarrow{\boldsymbol{a}} \text { have the same direction }\end{cases}$
: still holds if $\overrightarrow{\boldsymbol{F}}$ varies with time they will not, in general, be tangent to the path.

(a)

(b)

(c)

Fig. 12.1

### 12.1A Newton's $2^{\text {nd }}$ Law of Motion

Several forces

$$
\begin{equation*}
\sum \vec{F}=m \vec{a} \tag{12.2}
\end{equation*}
$$

$\uparrow$
sum, or resultant, of all the forces
Frame of reference
: system of axes with respect to which $\overrightarrow{\boldsymbol{a}}$ is determined is not arbitrary.
$\{$ must have a constant orientation with respect to the stars
$\{$ their origin must either be attached to the sun
or move with a constant velocity w. r. t. the sun
(sun $\leftarrow$ mass center of the solar system)
$\rightarrow$ Newtonian frame of reference (inertial system)
Precisely, axes attached to the earth $\rightarrow \mathrm{X}$ Newtonian frame of reference
However, in most engineering applications, enough for (12.1) and (12.2) without any applicable error.
If $\overrightarrow{\boldsymbol{a}}$ represents a relative acceleration w. r. t. moving axes, (12.1)
do not hold. (e.g. attached to an accelerated car rotating piece of machinery)
$\sum \overrightarrow{\boldsymbol{F}}=\mathbf{0} \rightarrow$ if initially at rest ( $\overrightarrow{\boldsymbol{v}}_{0}=\mathbf{0}$ ), remain at rest originally moving with $\vec{v}_{0}$, maintain constant velocity (in a straight line)
$\rightarrow$ Newton's $1^{\text {st }}$ law : particular case of the $2^{\text {nd }}$ law.

### 12.1B Linear Momentum, Rate of change of Linear Momentum

$$
\vec{a}=\frac{d \vec{v}}{d t} \quad, \quad \sum \vec{F}=m \frac{d \vec{v}}{d t}
$$

$m$ is constant,

$$
\begin{equation*}
\sum \vec{F}=\frac{d}{d t}(m \vec{v}) \tag{12.3}
\end{equation*}
$$

$\boldsymbol{m} \overrightarrow{\boldsymbol{v}}$ : (linear) momentum, same direction with velocity
(12.3) : the resultant of the forces acting on the particle is equal to the rate of change of the linear momentum of the particle.
(Newton's original statement regarding $2^{\text {nd }}$ law)
$\vec{L}$ : linear momentum

$$
\begin{align*}
& \vec{L}=m \vec{v} \\
& \sum \vec{F}=\dot{\vec{L}} \tag{12.5}
\end{align*}
$$

mass is assumed to be constant

### 12.1B Linear Momentum, Rate of change of Linear Momentum

(12.3), (12.5) $\rightarrow$ not applicable to the rockets, which gain or lose mass However, applicable to relativistic mechanics, where mass is assumed to vary with the particle speed

If $\sum \overrightarrow{\boldsymbol{F}}=\mathbf{0}$, rate of change of $\boldsymbol{m} \overrightarrow{\boldsymbol{v}}$ is zero
$\rightarrow$ Principle of conservation of linear momentum (alternative statement of Newton's $1^{\text {st }}$ law)

### 12.1C System of Units

$\overrightarrow{\boldsymbol{F}}=\boldsymbol{m} \overrightarrow{\boldsymbol{a}}$, units of force, mass, length and time cannot be chosen arbitrary
can choose three of four units arbitrarily but must choose the fourth unit
so that $\overrightarrow{\boldsymbol{F}}=\boldsymbol{m} \overrightarrow{\boldsymbol{a}}$ is satisfied.
$\rightarrow$ coherent set of kinetic units

International System of Units (SI Units)

force $\rightarrow$ derived unit, newton ( N )

$$
1 N=(1 \mathrm{~kg})\left(1 \mathrm{~m} / \mathrm{s}^{2}\right)=1 \mathrm{~kg} \cdot \mathrm{~m} / \mathrm{s}^{2}
$$

$$
\mathbf{a}=1 \mathrm{~m} / \mathrm{s}^{2}
$$

$$
m=1 \mathrm{~kg}
$$

$$
\mathbf{F}=1 \mathrm{~N}
$$

Fig. 4 A force of 1 newton gives a 1kilogram mass an acceleration of $1 \mathrm{~m} / \mathrm{s}^{2}$

SI base units: chosen and defined to be independent of the location, can be used anywhere on earth

### 12.1C System of Units

Weight $\vec{W}: W=m g$

$$
W=(1 \mathrm{~kg})\left(9.81 \mathrm{~m} / \mathrm{s}^{2}\right)=9.81 \mathrm{~N}
$$

Multiple/submultiples of units

$$
\begin{array}{ll}
1 \mathrm{~km}=1000 \mathrm{~m}, & 1 \mathrm{~mm}=0.001 \mathrm{~m} \\
1 \mathrm{Mg}=1000 \mathrm{~kg}=1 t, & 1 g=0.001 \mathrm{~kg} \\
1 \mathrm{kN}=1000 \mathrm{~N} &
\end{array}
$$

Unit of linear momentum

$$
m v=(k g)(m / s)=k g \cdot m / s
$$

U.S. customary units: foot (ft), pound (lb), second (s)

$$
1 f t=0.3048 m, \quad 1 l b=0.4536 k g f
$$

## $m=1 \mathrm{~kg}$

$\mathbf{W}=9.81 \mathrm{~N}$

Fig. 5 In the SI system, a block with mass 1 kg has a weight of 9.81 N


Fig. 6 In the U.S. system, a block with a weight of 1 lb in free fall has an acceleration of $32.2 \mathrm{ft} / \mathrm{s}^{2}$

### 12.1C System of Units

Unit of mass consistent with ft, lb, s

$$
\begin{aligned}
& f=m a, \quad \quad 1 l b=(1 \text { slug })\left(1 f t / s^{2}\right) \\
& 1 \text { slug }=\frac{1 l b}{1 f t / s^{2}}=1 l b \cdot s^{2} / f t
\end{aligned}
$$

Slug is a mass 32.2 times larger than the mass of the standard pound.

$$
m=\frac{W}{g}, m v=(s l u g)(f t / s)=\left(l b \cdot s^{2} / f t\right)(f t / s)=l b \cdot s
$$

Conversion from one system of units to another

$$
\begin{aligned}
& 1 \mathrm{ft}=0.3048 \mathrm{~m}, \quad 1 \mathrm{lb}=4.448 \mathrm{~N} \\
& 1 \mathrm{slug}=1 \mathrm{lb} \cdot \mathrm{~s}^{2} / f t=14.59 \mathrm{~kg} \\
& 1 \text { pound }- \text { mass }=0.4536 \mathrm{~kg}
\end{aligned}
$$

$$
\mathrm{a}=1 \mathrm{ft} / \mathrm{s}^{2}
$$

$\mathbf{F}=1 \mathrm{lb}$
( $=1 \mathrm{lb} \cdot \mathrm{s}^{2} / \mathrm{ft}$ )

Fig. 7 In the U.S. customary system, a force of 1 lb applied to a block with a mass of 1 slug produces an acceleration of $1 \mathrm{ft} / \mathrm{s}^{2}$

### 12.1D Eqn. of Motion

$$
\vec{F}=m \vec{a}
$$

- 2 important graphical tools to solve the dynamics problem using Newton's $2^{\text {nd }}$ law: free-body diagram (FBD), kinematic diagram (KD)

FBD: - body: multiple diagrams when necessary

- axes: Cartesin, normal/tangential, radial/transverse
- support forces: 2 perpendicular forces for a pin, normal force, friction force
- applied forces and body forces: weight, magnetic forces, known pulling force
- dimension: angle, distance

KD: - body: same in FBD

- inertial term: $\boldsymbol{m} \overrightarrow{\boldsymbol{a}}$ consistent with the coord. System or components


Fig. 9 Steps in drawing a free-body diagram and a kinetic diagram for solving dynamics problems

### 12.1D Eqn. of Motion

More convenient to replace $\overrightarrow{\boldsymbol{F}}=\boldsymbol{m} \overrightarrow{\boldsymbol{a}}$ by equivalent equations Involving scalar quantities
i) Rectangular components
: resolving $\overrightarrow{\boldsymbol{F}}$ and $\overrightarrow{\boldsymbol{a}}$ into rectangular components

$$
\begin{array}{rlrl}
\sum\left(F_{x} \vec{i}+F_{y} \vec{j}+F_{z} \vec{k}\right) & =m\left(a_{x} \vec{i}+a_{y} \vec{j}+a_{z} \vec{k}\right) \\
\sum F_{x}=m a_{x} & \sum F_{y} & =m a_{y} & \sum F_{z} \\
=m a_{z} \\
=m \ddot{x} & & =m \ddot{y} & \\
=m \ddot{z}
\end{array}
$$

: example of projectile $\vec{w}=-w \vec{j}$

$$
\begin{array}{ll}
m \ddot{x}=0, & m \ddot{y}=-w, \\
\ddot{x}=0, & m \ddot{z}=0 \\
\ddot{y}=-\frac{w}{m}=-g, & \ddot{z}=0
\end{array}
$$

$\longrightarrow$ can be integrated independently
: two or more bodies, $\rightarrow$ eqn. of motion should be written for each body all accelerations should be measured w. r. t. a Newtonian frame of reference

### 12.1D Eqn. of Motion

ii) Tangential and Normal Components

$$
\begin{align*}
\sum F_{t} & =m a_{t}, & \sum F_{n} & =m a_{n} \\
& =m \frac{d v}{d t} & & =m \frac{v^{2}}{\rho} \tag{12.9}
\end{align*}
$$

solve for two unknowns


Fig. 10 The net force acting on a particle moving in a curvilinear path can be resolved into components tangent to the path and normal to the path, producing tangential and normal components of acceleration

### 12.1D Eqn. of Motion

iii) Radial and Transverse Components

$$
\begin{align*}
& \sum F_{r}=m a_{r}, \quad \sum F_{\theta} \\
&=m a_{\theta}  \tag{12.10-12}\\
&=m\left(\ddot{r}-r \dot{\theta}^{2}\right) \quad=m(r \ddot{\theta}+2 \dot{r} \dot{\theta}) \\
& \longrightarrow \text { solve for two unknowns }
\end{align*}
$$



Fig. 12.14

### 12.1 Dynamic Equilibrium

Alternative form of Newton's $2^{\text {nd }}$ Law

$$
\begin{equation*}
\sum \vec{F}-m \vec{a}=0 \tag{12.10}
\end{equation*}
$$

If we add the vector $\mathbf{- m} \boldsymbol{m}$, we obtain a system of vectors equivalent to zero (Fig. 12.8)


Fig. 12.8
$-\boldsymbol{m} \overrightarrow{\boldsymbol{a}}$ : inertia vector
$=0$
Dynamic equilibrium : equilibrium under the given forces and inertia vector
, Closed-vector polygon : coplanar forces
$\sum \boldsymbol{F}_{x}=\mathbf{0}, \quad \sum \boldsymbol{F}_{y}=\mathbf{0}$ including inertia vector
> tangential and normal components :

$$
\text { inertia vector }\left\{\begin{array}{l}
-\boldsymbol{m} \vec{a}_{t} \\
-\boldsymbol{m} \vec{a}_{n}
\end{array}\right.
$$

### 12.1 Dynamic Equilibrium

Tangential component : measure of the resistance to a change in speed
Normal component : tendency to leave its curved path(centrifugal force)

Either $=0$ under special conditions
i) Start from rest, initial velocity $=0 \rightarrow$ normal component of inertia vector $=0$
ii) Constant speed $\rightarrow$ tangential component $=0$

Inertia vectors is often called "inertia forces" : measure of resistance when we set them in a motion or we try to change the conditions of motion Inertia forces $\neq$ forces found in statics (e.g. contact forces, gravitational forces)

### 12.2A Angular Momentum

Angular momentum, moment of momentum : moment about O of the vector $\boldsymbol{m} \overrightarrow{\boldsymbol{v}} \rightarrow \overrightarrow{\boldsymbol{H}}_{\text {。 }}$

$$
\vec{H}_{o}=\vec{r} \times m \vec{v}
$$


$\rightarrow$ perpendicular to the plane containing $\overrightarrow{\boldsymbol{r}}$ and $\boldsymbol{m} \overrightarrow{\boldsymbol{v}}$

$$
\left|\vec{H}_{o}\right|=r m v \sin \phi
$$

$\rightarrow$ angle between $\overrightarrow{\boldsymbol{r}}$ and $\boldsymbol{m} \overrightarrow{\boldsymbol{v}}$

$$
\text { unit : }(m)(k g \cdot m / s)=k g \cdot m^{2} / s
$$

Fig. 12.12

Resolving into component

$$
\overrightarrow{\boldsymbol{H}}_{o}=\left|\begin{array}{ccc}
\vec{i} & \vec{j} & \vec{k} \\
\boldsymbol{x} & \boldsymbol{y} & z \\
\boldsymbol{m} \boldsymbol{v}_{x} & \boldsymbol{m} \boldsymbol{v}_{y} & \boldsymbol{m} v_{z}
\end{array}\right| \quad \begin{aligned}
& \boldsymbol{H}_{x}=\boldsymbol{m}\left(\boldsymbol{y} \boldsymbol{v}_{x}-\boldsymbol{z} \boldsymbol{v}_{y}\right) \\
& \boldsymbol{H}_{y}=\boldsymbol{m}\left(\boldsymbol{z} \boldsymbol{v}_{x}-\boldsymbol{x} \boldsymbol{v}_{z}\right) \\
& \boldsymbol{H}_{z}=\boldsymbol{m}\left(\boldsymbol{x} v_{y}-\boldsymbol{y} \boldsymbol{v}_{x}\right)
\end{aligned}
$$

### 12.2A Angular Momentum

In case of a particle moving in the $x y$-plane

$$
\begin{aligned}
z=\boldsymbol{v}_{z}=\mathbf{0}, \quad \boldsymbol{H}_{y}=\boldsymbol{H}_{z}=\mathbf{0} \\
\boldsymbol{H}_{0}=\boldsymbol{H}_{z}=\boldsymbol{m}\left(\boldsymbol{x} \boldsymbol{v}_{y}-\boldsymbol{y} \boldsymbol{v}_{x}\right) \\
\quad \rightarrow \text { perpendicular to the xy-plane }
\end{aligned}
$$

$(+)$ or (-) according to the sense in which the particle is observed to move from $\boldsymbol{O}$
Polar coordinate : $\boldsymbol{H}_{0}=\boldsymbol{r m v} \boldsymbol{\operatorname { s i n }} \boldsymbol{\phi}=\boldsymbol{r m} \boldsymbol{v}_{\theta}$

$$
\begin{equation*}
=m r^{2} \dot{\theta}\left(v_{\theta}=r \dot{\theta}\right) \tag{12.18}
\end{equation*}
$$

Derivative w. r. t. $t$


Fig. 12.13

$$
\dot{\vec{H}}_{\mathbf{0}}=\dot{\vec{r}} \times \boldsymbol{m} \overrightarrow{\boldsymbol{v}}+\overrightarrow{\boldsymbol{r}} \times \boldsymbol{m} \dot{\vec{v}}=\underbrace{\overrightarrow{\mathbf{v}}}_{\substack{\downarrow \\ \mathbf{v}} \boldsymbol{m} \overrightarrow{\boldsymbol{v}}}+\overrightarrow{\boldsymbol{r}} \times \boldsymbol{m} \overrightarrow{\boldsymbol{a}} \boldsymbol{\downarrow}
$$

$\therefore \vec{v}$ and $m \vec{v}$
are collinear
$\dot{\overrightarrow{\boldsymbol{H}}}_{\mathrm{o}}=\dot{\vec{r}} \times \sum \overrightarrow{\boldsymbol{F}}=\sum \overrightarrow{\boldsymbol{M}}_{\mathrm{o}}$
: sum of the moments about $\boldsymbol{O}$ of the force
$=$ rate of change of angular momentum about $\boldsymbol{O}$

### 12.1D Eqn. of Motion in terms of Radial and Transverse Components

Polar coordinate $\boldsymbol{r}, \boldsymbol{\theta}$

$$
\begin{aligned}
& \sum F_{r}=m a_{r}, \quad \sum F_{\theta}=m a_{\theta} \\
& =m(\ddot{r}-r \ddot{\theta}) \quad=m(r \ddot{\theta}+2 \dot{r} \dot{\theta}) \\
& \text { or recalling } \quad \sum \overrightarrow{\boldsymbol{M}}_{0}=\dot{\overrightarrow{\boldsymbol{H}}}_{0}, \quad \boldsymbol{H}_{0}=\boldsymbol{m r ^ { 2 }} \dot{\boldsymbol{\theta}} \\
& \sum M_{0}=r \sum F_{\theta} \\
& r \sum F_{\theta}=\frac{d}{d t}\left(m r^{2} \dot{\theta}\right) \\
& =m\left(r^{2} \ddot{\theta}+2 r \dot{r} \dot{\theta}\right) \\
& \sum F_{\theta}=m(r \ddot{\theta}+2 \dot{r} \dot{\theta})
\end{aligned}
$$



Fig. 12.14

### 12.2B Motion under a Central Force

Moving under a central force : force $\overrightarrow{\boldsymbol{F}}$ directed toward or away from

$$
\begin{align*}
& \text { a fixed point } \mathrm{O} \text { (center of force) } \\
& \rightarrow \quad \sum \overrightarrow{\boldsymbol{M}}_{0}=\mathbf{0} \\
& \rightarrow \quad \dot{\overrightarrow{\boldsymbol{H}}}_{0}=\mathbf{0}, \quad \overrightarrow{\boldsymbol{H}}_{0}=\text { const } \tag{12.23}
\end{align*}
$$

Angular momentum of a particle moving under a central force is constant, in both magnitude and direction

$$
\rightarrow \quad \vec{r} \times m \vec{v}=\vec{H}_{0}=\text { const }
$$

$\rightarrow \overrightarrow{\boldsymbol{r}}$ must be perpendicular to the constant vector $\overrightarrow{\boldsymbol{H}}_{0}$, moves in a fixed plane Perpendicular to $\overrightarrow{\boldsymbol{H}}_{0}$
$\overrightarrow{\boldsymbol{H}}_{0}$ and the fixed plane are defined by the initial vector $\overrightarrow{\boldsymbol{r}}_{0}$ and $\overrightarrow{\boldsymbol{v}}_{0}$
Since $\left|\overrightarrow{\boldsymbol{H}}_{0}\right|$ is constant, $r m v \sin \phi=r_{0} m v_{0} \sin \phi_{0}$
Application : planetary motion, space vehicles in orbit about the earth

Fig. 12.15


Fig. 12.16

### 12.2B Motion under a Central Force

Polar coordinate

$$
m r^{2} \dot{\theta}=H_{0}=\text { const } .
$$

Divide by $\boldsymbol{m}, \boldsymbol{h}$ : angular momentum per unit mass $\boldsymbol{H}_{0} / \boldsymbol{m}$

$$
\begin{equation*}
\boldsymbol{r}^{2} \dot{\boldsymbol{\theta}}=\boldsymbol{h} \tag{12.25}
\end{equation*}
$$

Fig. 12. 15, radius vector OP sweeps an infinitesimal area


Fig. 12.15
> Moving under a central force, areal velocity is constant.

### 12.2C Newton's Law of Gravitation

Example of central force : $\{$ force exerted by the sun or the planet force exerted by the earth or orbiting satellite
$\rightarrow$ how to determine the magnitude of gravitational force
Law of universal Gravitation : two particles of masses $\boldsymbol{M}$ and $\boldsymbol{m}$ at a distance $\boldsymbol{r}$ attract each other with equal and opposite forces $\overrightarrow{\boldsymbol{F}}$ and $-\overrightarrow{\boldsymbol{F}}$ directed along the line joining the particles.

$$
F=G \frac{M m}{r^{2}}
$$

$\boldsymbol{G}$ : universal constant, "constant of gravitation"
$(66.73 \pm 0.03) \times 10^{-12} \mathrm{~m}^{3} / \mathrm{kg} \cdot \mathrm{s}^{2}$
>Affect becomes appreciable only when one of the bodies has a very large mass
[ a planet about the sun
\{ satellite orbiting the earth
bodies falling on the surface of the earth

### 12.2C Newton's Law of Gravitation

Weight $\vec{W}$ : force exerted by the earth on a body located on or near its surface

$$
W=m g=\frac{G M}{R^{2}} m \quad \text { or } \quad g=\frac{G M}{R^{2}}
$$

Earth is not truly spherical $\rightarrow$ the value of $\boldsymbol{W}$ and $\boldsymbol{g}$ vary with the altitude and latitude
Reason \#2 of varying $\boldsymbol{W}$ and $\boldsymbol{g}$ : system of axes attached to the earth is NOT a Newtonian frame of reference
> include centrifugal force due to the earth rotation $g \simeq 9.781 \mathrm{~m} / \mathrm{s}^{2}$ (at the equator), $g \simeq 9.833 \mathrm{~m} / s^{2}$ (at the poles)
> $g=9.7807\left(1+0.0053 \sin ^{2} \phi\right) m / s^{2}, \phi$ : latitude (위도) longitude (경도)
Force exerted by the earth on a body of mass $\boldsymbol{m}$ located in space at a distance $\boldsymbol{r}$

$$
F=G \frac{M m}{r^{2}}
$$

Can be simplified by $\boldsymbol{G M}=\boldsymbol{g} \boldsymbol{R}^{\mathbf{2}}$

$$
\left.\begin{array}{l}
g: 9.81 m / s^{2} \\
R: 6.37 \times 10^{6} m
\end{array}\right\} \leftarrow \text { average values }
$$

