

Ch. 15: Basic Concepts of Small-amplitude Waves

I-3

- * Relevant governing equations for plasma dynamics are non linear.
- * Assume amplitudes of waves are small enough (infinitesimally small).
 - ⇒ Equations can be linearized.
 - ⇒ Fourier analysis becomes useful.
- * 0-th order: equilibrium of the system.
- * 1-st order: Solve for frequency " ω " $\equiv \omega(\vec{k})$, and relative amplitudes and phases of various oscillating quantities.

* Exponential Notation:

For example, density perturbation:

$$\delta n = \underbrace{\delta \bar{n}}_{\substack{\text{amplitude} \\ \text{(real number)}}} \exp \left[i \left(\underbrace{\vec{k} \cdot \vec{x} - \omega t + \phi_n}_{\substack{\text{spatial, temporal} \\ \text{phase} \\ \text{dependences}}} \right) \right]$$

(cf: sometimes, $\delta \bar{n} e^{i\phi_n}$ is combined into a complex amplitude $\tilde{\delta n}$)

So, understanding the notations depending on the context of discussion is useful, rather than being trapped in irrelevant details.

Of course, $e^{iA} = \cos A + i \sin A$

is a useful identity.

Physical Quantities of Interests

I.-5.

* \vec{k} : wave vector

$$\lambda = 2\pi/k = 2\pi/|\vec{k}| \quad : \text{wavelength.}$$

* ω : frequency

$$\omega = \omega(\vec{k})$$

Dispersion Relation

* Phase Velocity :

"an observer travelling with the phase velocity, \vec{v}_p , stays at a constant wave phase."

⇒ Demand $\frac{d}{dt} (\text{total phase factor}) = 0$

$$\left(\text{in } \exp [i(\vec{k} \cdot \vec{x} - \omega t + \phi_0)] \right)$$

* Phase Velocity :

From $\frac{d}{dt} (\vec{k} \cdot \vec{x} - \omega t) = \vec{k} \cdot \frac{d\vec{x}}{dt} - \omega = 0,$

$$\vec{V}_p = \frac{\vec{k}}{|\vec{k}|^2} \omega = \frac{\omega k_x}{k^2} \hat{x} + \frac{\omega k_y}{k^2} \hat{y} + \frac{\omega k_z}{k^2} \hat{z}$$

in Cartesian coordinates



not

$$\frac{\omega}{k_x} \hat{x} + \frac{\omega}{k_y} \hat{y} + \frac{\omega}{k_z} \hat{z}$$

- * If we consider a wave-packet consisting of many components with different \vec{k} 's (and ω 's), \vec{V}_p is the velocity at which individual crests (within the packet) travel.

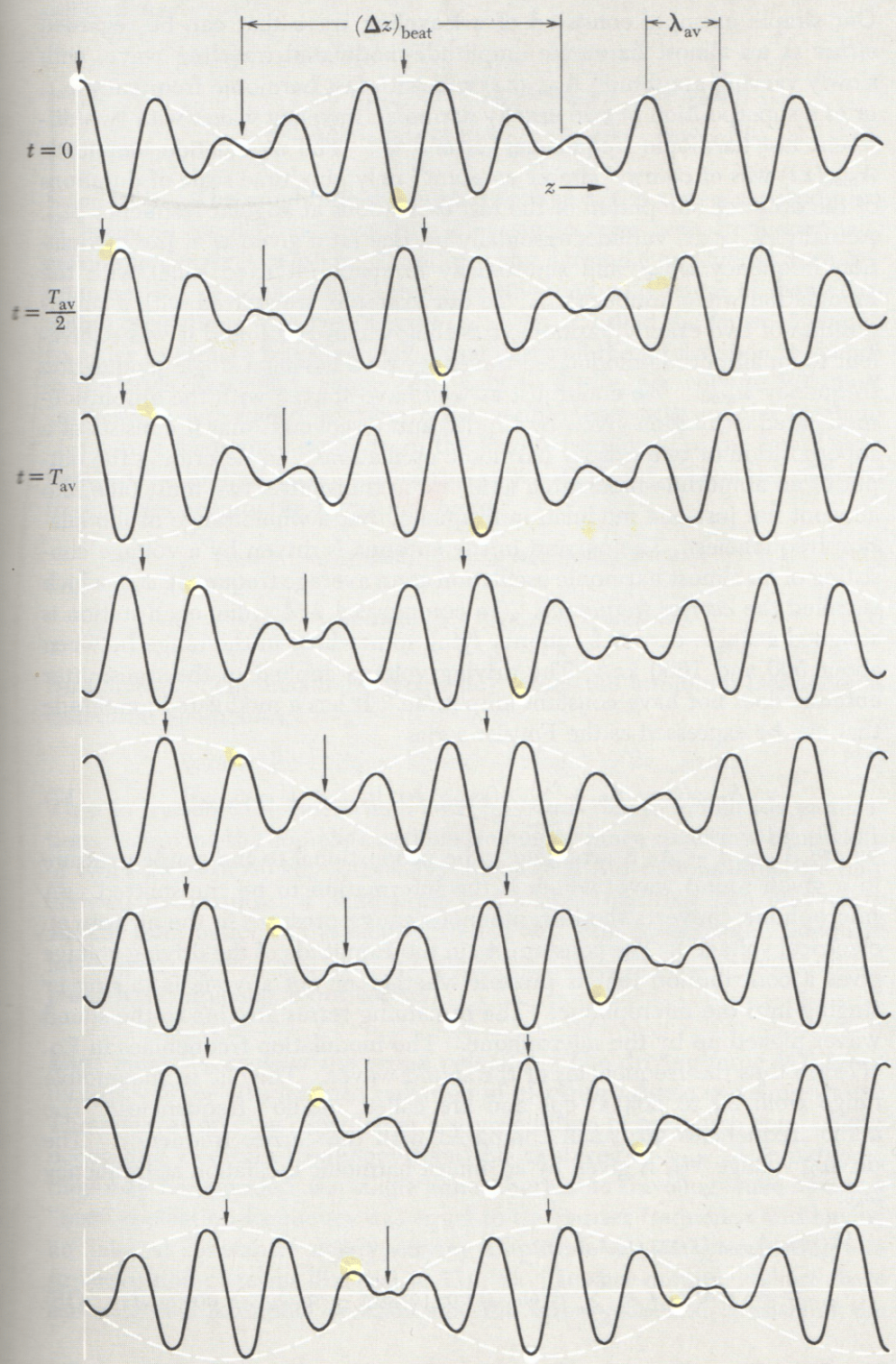


Fig. 6.1 Group velocity. The arrows follow the beats, which travel at the group velocity v_g . The white circles follow individual wave crests, which travel at the average phase velocity v_{av} .

draw these

1.-7.

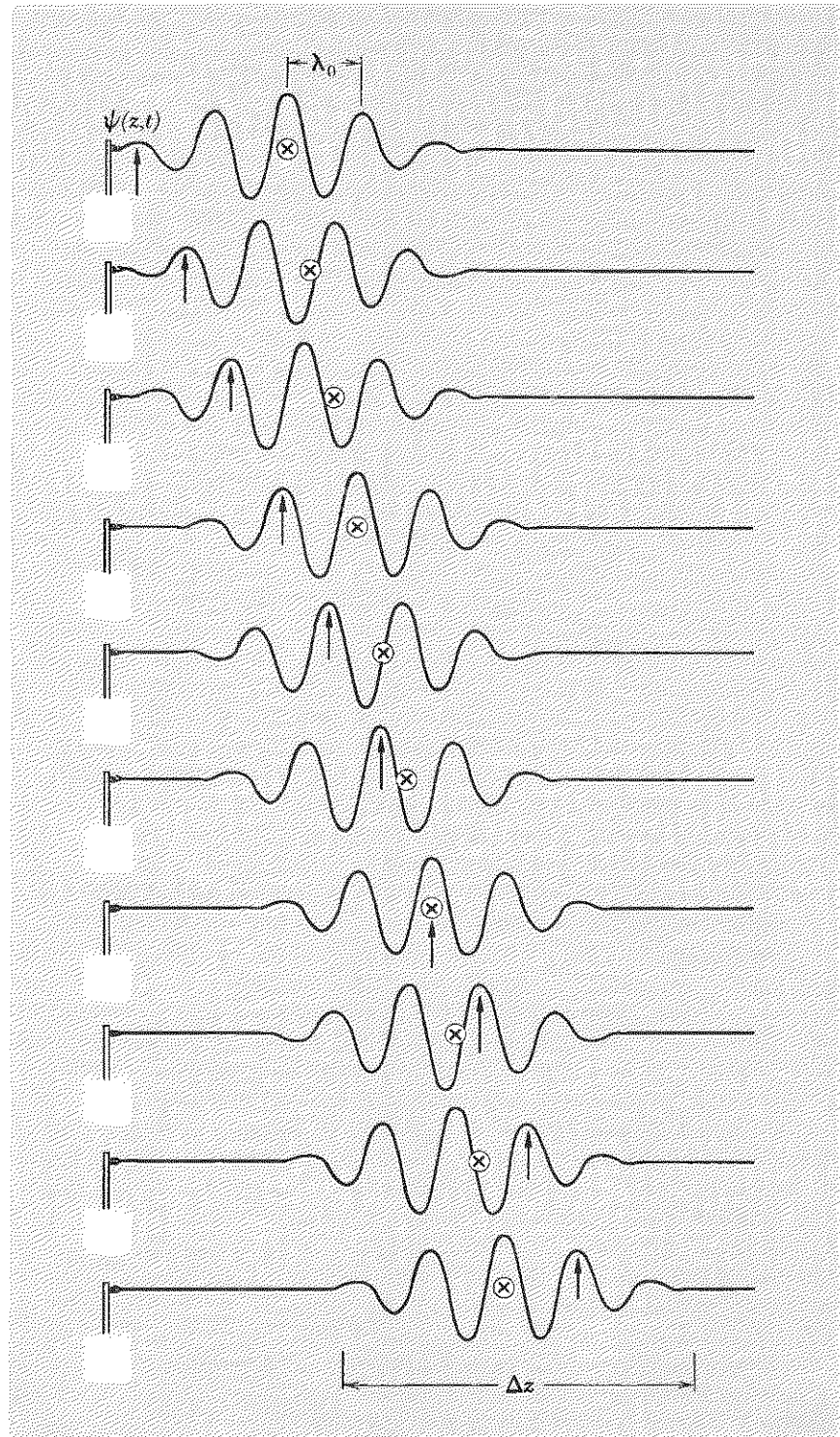
of envelope

of crests & troughs
 $\frac{1}{4} \text{ wl.}$

then packet prop. by

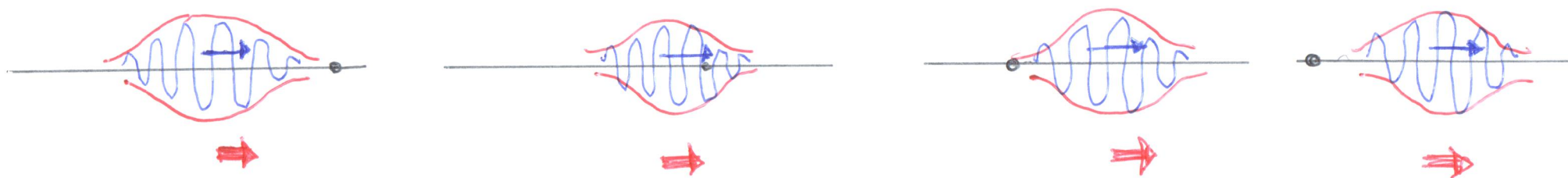
$\frac{1}{4} \text{ wl.}$

Fig. 6.7 Wave packet with phase velocity twice the group velocity. The arrow travels at the phase velocity, following a point of constant phase for the dominant wavelength. The cross travels at the group velocity with the packet as a whole.



15.2. Group Velocities

I.-8.



- * \vec{v}_p : phase velocity of individual crests within the packet,
(cf. it can be in opposite direction to \Rightarrow of packet)
- * \vec{v}_g : group velocity of the wave packet
(\sim envelope containing many wave crests).
- * In general, $\vec{v}_g \neq \vec{v}_p$.
- * Energy and Information are carried at \vec{v}_g !

* Packet of oscillations with Gaussian envelope;

$$\bullet A(x) = \operatorname{Re} \left[\exp\left(-\frac{x^2}{2\sigma^2}\right) \exp(ik_0 x) \right] \quad (15.8)$$

See Fig 15.1 in G&R ; $k_0 \sigma \gg 1$ (typo in Fig. caption)

It can be shown that ,

$$\bullet A(x) = \operatorname{Re} \left(\frac{\sigma}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dk \exp(ikx) \exp\left[-\frac{\sigma^2(k-k_0)^2}{2}\right] \right) \quad (15.9)$$

*** Home work: Problem 15.2. on page 253

This can be viewed as $t=0$ snapshot (free-frames) of a set of propagating waves.

* Ensuing time-evolution of this system is then,

$$A(x,t) = \text{Re} \left(\frac{\sigma}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dk \exp[i(kx - \omega(k)t)] e^{-\frac{\sigma^2(k-k_0)^2}{2}} \right) \quad (15.10)$$

Note:

- Dependences of Gaussian envelope on σ :
 - large $\sigma \rightarrow$ narrow (localized) packet in k-space

\rightarrow wide packet in x -space.

(and vice versa for small σ)

\sim uncertainty principle in quantum mechanics.

approximate;

$$\omega(k) = \omega(k_0) + \left(\frac{\partial \omega}{\partial k} \right)_{k_0} (k - k_0) + \dots$$

⇒

$$A(x,t) = \text{Re} \left[\exp \left[i \left\{ k_0 \left(\frac{\partial \omega}{\partial k} \right)_{k_0} - \omega(k_0) \right\} t \right] \right. \\ \left. \times \int_{-\infty}^{\infty} dk \exp \left\{ i \left[\underbrace{kx - k \left(\frac{\partial \omega}{\partial k} \right)_{k_0} t}_{\text{phase}} \right] \right\} e^{-\frac{\sigma^2 (k-k_0)^2}{2}} \right]$$

(15.11)

Comparing this expression to $A(x, t=0)$ in Eq (15.9),

the second line " \sim " is exactly " $A(x - \left(\frac{\partial \omega}{\partial k} \right)_{k_0} t, 0)$ "

i.e., a translation of the original $t=0$ freeze-frame
at velocity $\left(\frac{\partial \omega}{\partial k} \right)_{k_0}$!

15.3. Ray-tracing Equations

I.-12.

* Consider an inhomogeneous plasma in 3d.

Without loss of generality, take $\vec{k}_0 = k_0 \hat{x}$.

Then, 3d-extension of Eq. (15.8) is:

$$A(x) = \text{Re} \left[\exp \left[-\frac{x^2}{2\sigma_x^2} - \frac{y^2}{2\sigma_y^2} - \frac{z^2}{2\sigma_z^2} \right] \exp(ik_0 x) \right]$$

which can be re-expressed through the Fourier transform, (15.12)

$$A(x) = \text{Re} \left[\frac{\sigma_x \sigma_y \sigma_z}{(2\pi)^{3/2}} \int_{-\infty}^{\infty} d\vec{k} \exp(i\vec{k} \cdot \vec{x}) \right]$$

$$\exp \left[-\sigma_x^2 (k_x - k_0)^2 / 2 - \sigma_y^2 k_y^2 / 2 - \sigma_z^2 k_z^2 / 2 \right]$$

(15.13)

* As done in 1d before, expand

I.-13.

$$\omega(\vec{k}) \simeq \omega(\vec{k}_0) + (\vec{k} - \vec{k}_0) \cdot \vec{\nabla}_{\vec{k}} \omega \Big|_{\vec{k}_0} + \dots \quad (15.14)$$

where

$$\vec{\nabla}_{\vec{k}} \equiv \hat{x} \frac{\partial}{\partial k_x} + \hat{y} \frac{\partial}{\partial k_y} + \hat{z} \frac{\partial}{\partial k_z} \quad (15.15)$$

* A similar analysis as Eqs (15.9) - (15.11) in 3d will lead to our 'freeze-frame' $\star (\vec{x})$ translating at a vector **group velocity** given by $(\star (\vec{x} - \vec{v}_g t))$,

$$\vec{v}_g = \frac{\partial \omega}{\partial \vec{k}} \quad (= \vec{\nabla}_{\vec{k}} \omega)$$

* Note that \vec{v}_g may not only have a different magnitude from \vec{v}_p , but even a different direction,

* As a packet propagates in an inhomogeneous plasma, the peak of \vec{k} spectrum, \vec{k}_0 can change.

* On the other hand, if the \checkmark medium is by hypothesis (background), linear and **time-independent**, $\omega(\vec{k}_0)$ should be constant.

\Rightarrow Total derivative of ω , moving with the wave-packet must vanish. i.e.,

$$\frac{d}{dt} \omega(\vec{k}, \vec{x}) = 0$$

along the ray.

* Using the chain-rule,

$$\frac{d}{dt} \omega(\vec{k}, \vec{x}) = \frac{d\vec{x}_0}{dt} \cdot \frac{\partial}{\partial \vec{x}} \Big|_{\vec{k}} \omega + \frac{d\vec{k}_0}{dt} \cdot \frac{\partial}{\partial \vec{k}} \Big|_{\vec{x}} \omega = 0$$

∴ we can identify the relation; (15.17)

- $\frac{d\vec{k}_0}{dt} = - \frac{\partial \omega}{\partial \vec{x}} \Big|_{\vec{k}}$
- $\frac{d\vec{x}_0}{dt} = \frac{\partial \omega}{\partial \vec{k}} \Big|_{\vec{x}}$

(15.18)

* Ray-tracing equations for wave-packet -

* As the wave-packet propagates it maintains the peak of its frequency spectrum, while \vec{k} -spectrum transforms.

* Ray-tracing Equation

for Wave-Packet : \leftarrow Wave-Particle \rightarrow
duality

Hamiltonian Mechanics
for Particle

$$\frac{d\vec{k}_0}{dt} = - \frac{\partial}{\partial \vec{x}} \omega$$

$$\frac{d\vec{x}_0}{dt} = \frac{\partial}{\partial \vec{k}} \omega$$

$$\hbar \omega = E$$

$$\hbar \vec{k} = \vec{p}$$

in Q.M.

$$\frac{d\vec{p}}{dt} = - \frac{\partial}{\partial \vec{q}} H(\vec{p}, \vec{q})$$

$$\frac{d\vec{q}}{dt} = \frac{\partial}{\partial \vec{p}} H(\vec{p}, \vec{q})$$

* \vec{v}_g of de Broglie wave \sim \vec{v} particle.

* Non conservation of \vec{k}_0 in an inhomogeneous medium
is related to the non-conservation of momentum

* Conservation of ω in time-independent medium
is related to the conservation of energy.

Ch. 16. Waves in an unmagnetized plasma

II-1

16.1. Langmuir Waves

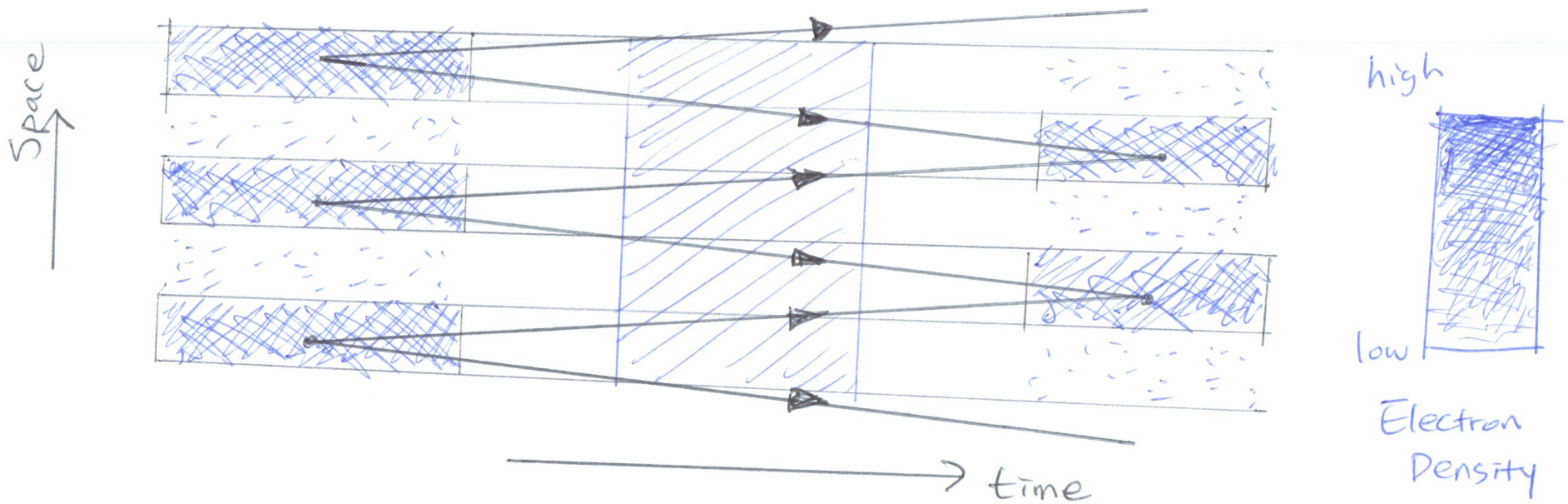


Fig 16.1.

- * Initial displacements of electrons from equilibrium position
 - charge separation → electric field generation
 - pulling electrons back → overshoot → oscillations,

* This oscillation happens in very short time scale

⇒ ions with heavy mass cannot respond,
and can be considered as stationary.

* Electron Dynamics :

$$- m n_e [\dot{\vec{u}}_e + (\vec{u}_e \cdot \vec{\nabla}) \vec{u}_e] = - e n_e \vec{E} - \vec{\nabla} p_e \quad (16.1)$$

$$- \dot{n}_e + \vec{\nabla} \cdot (n_e \vec{u}_e) = 0 \quad (16.2)$$

$$- \epsilon_0 \vec{\nabla} \cdot \vec{E} = e (n_i - n_e) \quad (16.3)$$

* Consider electrons moving in x direction and
waves propagating " " .

⇒ perturbed quantities vary as "exp [i (kx - ωt)]"

$$\frac{\partial}{\partial x} \rightarrow ik \quad \text{and} \quad \frac{\partial}{\partial t} = -i\omega$$

* Linearize Eqs. (16.1)-(16.3) i.e., keep only the 1st order quantities:

In addition, relate p_1 to n_1 by using the equation of state for 1-dimensional compression faster than the thermal conduction, i.e.,

$$p \propto n^\gamma, \text{ where } \gamma = \frac{2+N}{N} = \frac{2+1}{1} = 3.$$

$$\Rightarrow \left[\begin{array}{l} i\omega m n_0 u_1 = e n_0 E_1 + 3 i k T n_1 \end{array} \right. \quad (16.6)$$

$$\left[\begin{array}{l} -i\omega n_1 + i k n_0 u_1 = 0. \end{array} \right. \quad (16.7)$$

$$\left[\begin{array}{l} i k \epsilon_0 E_1 = -e n_1 \end{array} \right. \quad (16.8)$$

3 equations for 3 perturbed quantities, u_1 , n_1 and E_1 .

* [Determinant] = 0
of 3x3 matrix.

\Rightarrow Dispersion Relation
(between " ω " and " k ")

*

$$\omega^2 = \omega_{pe}^2 + 3k^2 v_{Te}^2$$

where

$$\omega_{pe}^2 \equiv n_e e^2 / \epsilon_0 m_e$$

electron plasma frequency

$$v_{Te}^2 = (T_e / m_e)$$

electron thermal velocity.

Bohm-Gross dispersion relation [PR 1949]

* Fig 16.2 illustrates the behavior at different k regimes.

⊗ At long wavelength (low k); $\omega^2 \approx \omega_{pe}^2$

\sim non-propagating "plasma oscillation"
(Langmuir)

⊗ At short wavelength (high k);

$$\omega^2 \approx 3k^2 v_{Te}^2$$

\sim Propagating "electron sound wave."

16.2. Ion Sound Waves

II-5.

- * Take a look at another electrostatic longitudinal wave ($\vec{k} \parallel \vec{E}_1$) which oscillates in longer time scale so that ions dynamics plays an essential role.
 - Electrons can move fast enough to establish nearly exact force balance (between pressure gradient and electric field), i.e., a Boltzmann response.
- * Follow a similar derivation as the one for Langmuir wave.
 - Ion Fluid momentum eqn. (16.16)
 - Poisson eqn. (16.19)
 - Ion continuity eqn (16.21)

* A crucial difference (from the Langmuir curve where $n_{i1} = 0$)

is that

$$n_{e1} = n_e - n_{e0} = n_{e0} (\exp(e\phi_1/T_e) - 1) \\ \approx n_{e0} \cdot e\phi_1/T_e \quad (16.18)$$

i.e. Boltzmann response (or adiabatic response),

* Linearization \rightarrow solve ~~the~~ coupled eqns for n_{i1} , ϕ_1 , and $U_{i1} \rightarrow$ to get:

$$\left(\frac{\omega}{k}\right)^2 = \frac{T_e/M}{1 + k^2 \lambda_D^2} + \gamma_i T_i / M \quad (16.23)$$

$$\lambda_D^2 \equiv \epsilon_0 T_e / n_e e^2 = v_{Te}^2 / \omega_p^2$$

Debye length.

* In the long wavelength limit (~~low~~ low k);

$$\star \quad \left(\frac{\omega}{k} \right)^2 \approx \frac{T_e + \gamma_i T_i}{M} \quad \text{Sound wave}$$

— Both electron and ion pressure contribute,
but only ions provide "mass".

— In collisionless plasma, Eq. \star is valid for $T_e \gg T_i$

because one needs to consider wave-particle
resonant ~~and~~ interaction (Landau damping) for

$\omega/k \sim v_{T_i}$ using kinetic theory.

* In the short wavelength limit (high k);

$$\omega^2 \approx \frac{T_e/M}{\lambda_D^2} \approx \left(\frac{m}{M} \right) \omega_p^2 \equiv \Omega_p^2; \quad \text{ion plasma frequency.}$$

non-propagating wave.

16.3. High-Frequency Electromagnetic Waves

II-8.

in an Unmagnetized Plasma.

* Now, we consider magnetic field perturbation \vec{B}_1 in addition to \vec{E}_1 . Then, we need to consider Ampère's law and Faraday's law. With $\vec{\nabla} \Rightarrow i\vec{k}$,

$$- i\vec{k} \times \vec{B}_1 = \mu_0 \vec{j}_1 - i\omega \vec{E}_1 / c^2 \quad (16.24)$$

$$- i\vec{k} \times \vec{E}_1 = i\omega \vec{B}_1 \quad (16.25)$$

⊗ Taking $\vec{k} \times$ Eq. (16.25) and expanding the triple vector product,

$$\boxed{k^2 \vec{E}_1 - \vec{k} (\vec{k} \cdot \vec{E}_1) = \left(\frac{\omega}{c}\right)^2 (\vec{E}_1 + i\vec{j}_1 / \epsilon_0 \omega)} \quad (16.27)$$

* For longitudinal waves with $\vec{k} \parallel \vec{E}_1$, LHS = 0 and we have an electrostatic wave in which Poisson eqn plays an important role.

* We consider here a transverse wave with $\vec{k} \perp \vec{E}_1$ (i.e., $\vec{k} \cdot \vec{E}_1 = 0$). \Rightarrow no need to consider Poisson eqn or density perturbation or continuity eqn.

* $\vec{J}_1 = -n_0 e \vec{u}_1$ (16.28)



where

(assume stationary ions for high frequency waves as before).

$-\bar{i}\omega m \vec{u}_1 = -e \vec{E}_1$ (16.29)

\Rightarrow Eq. (16.27) becomes,

$(c^2 k^2 - \omega^2) \vec{E}_1 = i\omega \vec{J}_1 / \epsilon_0 = -\left(\frac{n_0 e^2}{m \epsilon_0}\right) \vec{E}_1$ (16.31)

$$\omega^2 = c^2 k^2 + \omega_p^2 \quad (16.32)$$

Dispersion Relation for "EM wave" in unmagnetized plasma.

(*) Note that:

$$v_p \equiv \frac{\omega}{k} = c \left(1 + \frac{\omega_p^2}{c^2 k^2} \right)^{1/2} > c.$$

while

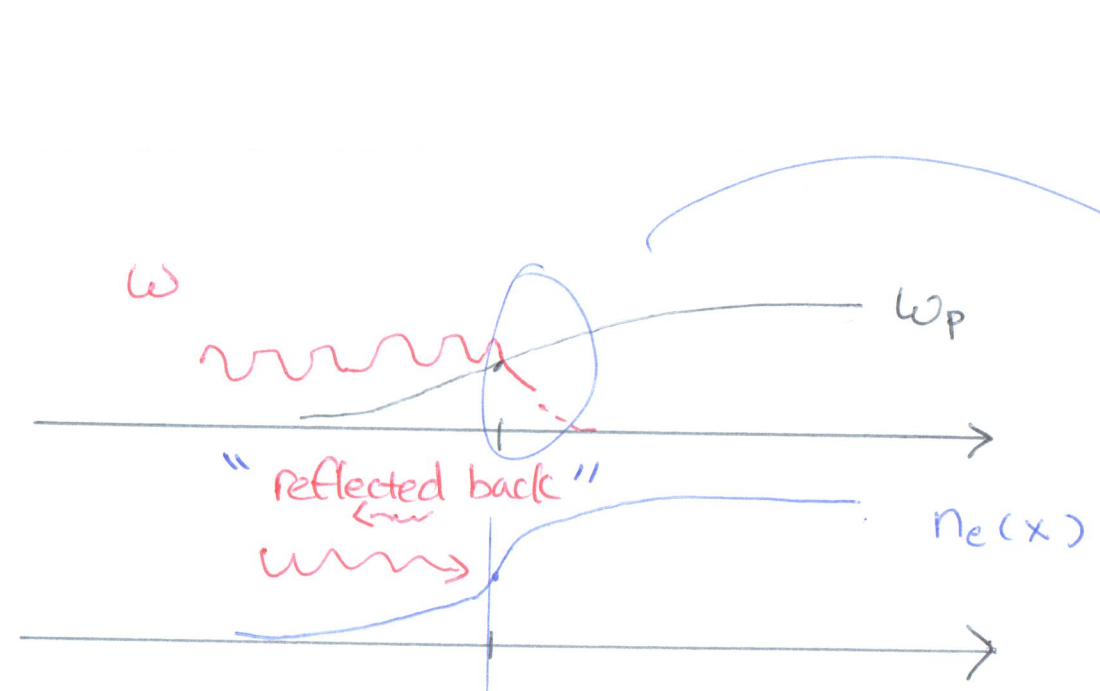
$$v_g \equiv \frac{\partial \omega}{\partial k} = \frac{c^2}{v_p} < c.$$

This is **NOT** a violation of the special relativity because the information and energy propagate with v_g **not** v_p .

- low k ; \rightsquigarrow constant freq. non-propagating ~~oscillation~~
- hi k ; \rightsquigarrow EM waves in vacuum propagating with "c".

(*) Homework Problem 16.2 on page 266.

(*) EM waves cannot propagate in a plasma with $\omega_p > \omega$.



"propagation" with $\omega > \omega_p$
 "propagation" not possible
 Cut-off layer
 with cut-off density

$$n_c = m_e \epsilon_0 \omega^2 / e^2.$$

For $\omega_p > \omega$,
 D.R. admits "k" which is imaginary,

i.e.,

$$k = (\omega^2 - \omega_p^2)^{1/2} / c$$

$$= \pm i (\omega_p^2 - \omega^2)^{1/2} / c$$

(16.37)

• $\exp(ikx) = \exp(-x (\omega_p^2 - \omega^2)^{1/2} / c)$
 evanescent solution

• Collisionless skin depth:
 $\equiv "c/\omega_p"$